

# European Extremely Large Telescope Astronomical instrumentation

21 September 2011

Wilfried Boland



# **NOVA** introduction

Netherlands Research School for Astronomy

 Top research school, evaluated exemplary in 2010
 Federation of university astronomy institutes
 290 fte scientific staff (20% is directly funded by NOVA)

 Mission

Facilitating top astronomical research in the Netherlands
 Hire researchers
 Build instruments

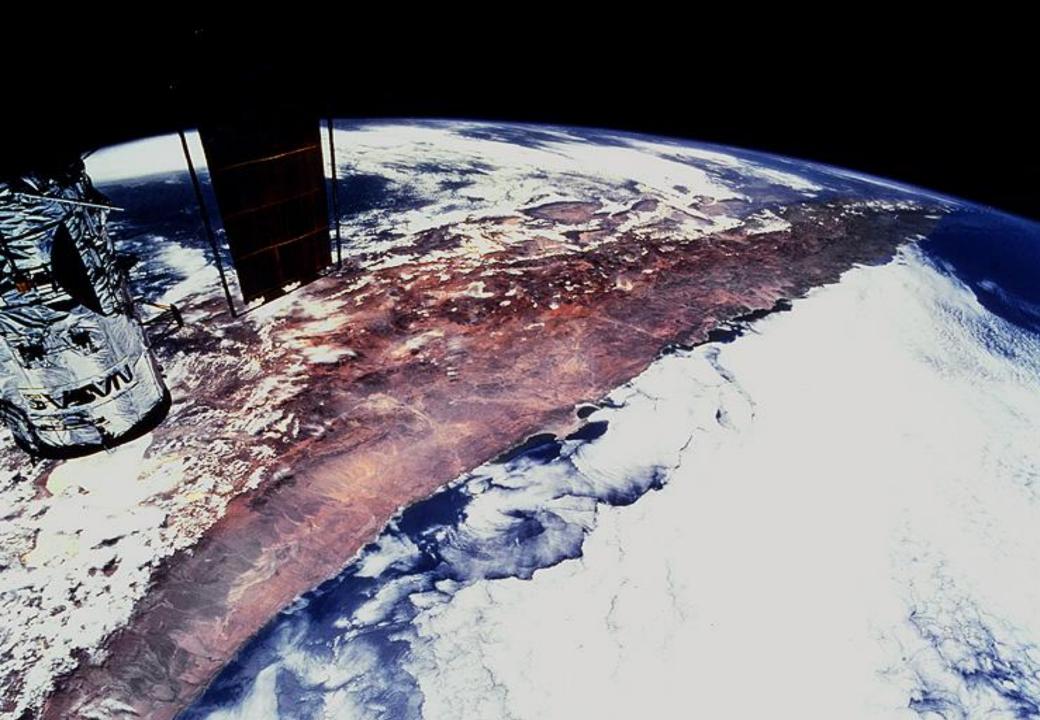
Train young astronomers at highest international level

# ESO Very Large Telescope

of the state

#### Atacama Large Millimeter Array ALMA

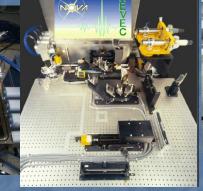




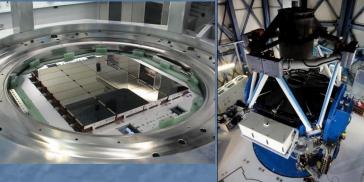
# NOVA ESO projecten



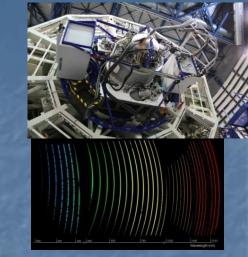
MIDI



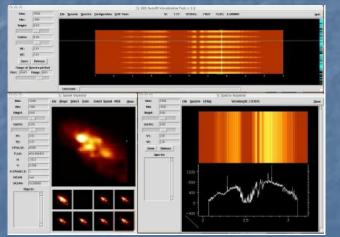
NEVEC



OmegaCAM voor VST  $\rightarrow$  OmegaCEN



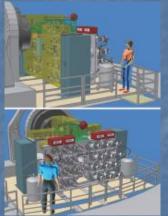
X-Shooter nabij-IR spectroscopische arm



SINFONI: 2k camera voor SPIFFI: nabij-IR integral field spectrometer



Optical bench voor SPHERE Zimpol



MUSE-ASSIST: test set-up voor nieuwe VLT deformeerbare secundaire spiegel

Nieuwe grote ESO project: ~40m optisch/IR telescoop Fase B afgerond: klaar voor de bouw! (na goedkeuring Council)



### NOVA instrumentation program

Many discoveries are driven by new instrument capabilities

- Involvement in instrument ⇒ decision about functionality
- Involvement in instrument ⇒ understanding the instrument performance
- Involvement in instrument ⇒ early access to data
- Involvement in instrument ⇒ ideal position to make discoveries!
- NOVA strategy:
  - Design & construct instruments for international facilities
    Focus on ESO
  - NOVA Optical-Infrared instrumentation group located at ASTRON in Dwingeloo

Astronomy in the Netherlands

Mid-term update of the Strategic-Plan 2001-2010 and forward look to 2015

#### **Instrument Project Characteristics**

- Collaborations with international partners (for ESO projects ~4-6 partners)
- NOVA astronomer NL-leader and connection to the international consortium
- Common Project management procedures under ESO protocol (PDR, FDR, progress meetings etc.)
- Hardware design and manufacturing by NOVA Optical-IR instrumentation group
- Dutch astronomers in (inter)national science team to ensure interesting capabilities



### NOVA and the E-ELT

#### Objective:

Participate in design & construction of instrumentation for E-ELT
 In one as a leading partner (40% share)
 In another one as minor partner (20% share)

#### Funding:

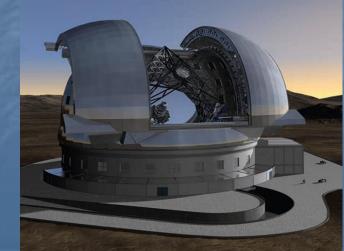
General NOVA budget

■ ESFRI grant of 18.78 M€

■ 8.78M€ for design and development

■ 10M€ to build one instrument (requires PI role)

Other grants



### E-ELT: 8 SCIENCE INSTRUMENTS +2 Post Focal AO MODULE STUDIES

INSTRUMENT MAIN OBSERVING MODES

OPTIMOS

CODEX

METIS

**EAGLE** 

HARMONI

SIMPLE

MICADO

EPICS + XAO

MAORYLTAO

Multi-slit and fiber MOS options are being studied High Resolution, High Stability Visual Spectrograph Mid IR camera /spectrograph WF, Multi IFU NIR Spectrograph. +AO Single IFU, Wide Spectral Band Spectrograph High-Resolution IR spectrograph NIR Camera sampling to the DF Planet Imager and Spectrograph

(MCAO) with 2 additional DM Module Provides DL images over a field <30"



#### **E-ELT Instrumentation in NL**

Consortium: Universities: NOVA, TU Delft, UTwente Technological institutes: ASTRON, SRON & TNO Companies: Airborne Composite BV, Dutch Space, JPE Applied for in 2008, awarded in 2009, end 2020+ Phase I (8.78M€): Preliminary design (4 instruments) Technology developments Phase II (10M€): Construction of one instrument

#### Industrial participation

Large research facilities means big business Industry can become project supplier to ESO and/or NOVA Construction of the telescope Delivery of subsystems to the telescope Supplier of parts of instrumentation, or partner in (optical, mechanical or thermal) design; partner is R&D to demonstrate technical readiness

### Industrial participation

Large research facilities means big business Industry can become project supplier to ESO and/or NOVA Construction of the telescope Delivery of subsystems to the telescope Supplier of parts of instrumentation, or partner in (optical, mechanical or thermal) design; partner is R&D to demonstrate technical readiness

#### Technology developments I

Vibration-free and precise cryo-coolers Present partners: UTwente, Dutch Space Motivation: High precision instruments, no vibrations Potential solution: sorption coolers <u>Remaining problems</u>: Sorption compresso Passive check valves • Cooling power too low (10mW  $\Rightarrow$ 1W) high pressure pressure buffer buffer ■ University product ⇒ commercial product

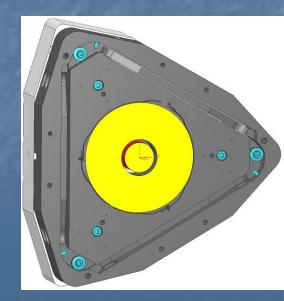
## Technology developments II

#### Movable cryogenic systems

- Present partners: NOVA Op-IR, JPE, SRON, TNO
- <u>Motivation</u>: High precision positioning and stability of movable elements in a cryogenic environment (80K)

#### Problems:

 Opto-mechanical engineering
 Very accurate positioning (nm), metrology and control



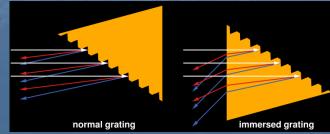
## Technology developments III

#### New optical components and materials

- <u>Present partners</u>: Airborne, NOVA Op-IR, SRON, TNO
- <u>Motivation</u>: Standard techniques will make the instrument rather big and heavy, or do not provide the required stability
- <u>Potential solutions</u>: composite materials, immersed gratings, integrated optics, smart optics, free form mirrors

#### <u>Remaining problems</u>:

- Behavior of composite materials in a cryo-vacuum environment (stiffness, air tightness, out-gassing)
- Immersed gratings have not yet been made with the required accuracy
- Manufacture products with the required accuracy (required micro-roughness RMS for free-form 30cm large Al mirrors = 15nm)



### Technology developments IV

Other areas where NOVA will look for partners: Polarimetric elements and engineering Precision engineering Better performance prediction, improved overall system engineering control, modeling alignment tolerances, Advanced data flow system AO Control (hardware and software) Industrial production process

#### There are many opportunities for industrial participation and products!

20

# One example: METIS

Mid-infrared E-ELT Imager and Spectrograph Operating from 3 to 14 micron Imager (L,M, N-band) Low resolution long slit spectrometer (L, M, N-band) High resolution IFU spectrometer (L, M band) Coronography (L, M, N band) Polarimetry (N-band) NOVA has PI role Overall project management High resolution IFU spectrograph Fore optics Cold central structure

#### Conclusion and Contact

There are many opportunities for industrial participation and products!

Wilfried Boland (<u>boland@strw.leidenuniv.nl</u>)
 Frank Molster (<u>molster@strw.leidenuniv.nl</u>)