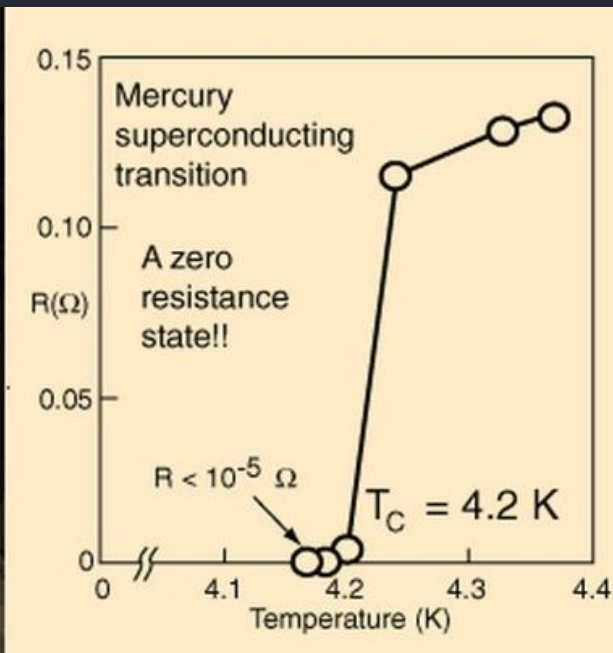


Superconducting high-resolution imaging X-ray spectrometers for space and ground based applications

Luciano Gottardi

Vision, Robotics & Motion, 's-Hertogenbosch, 8 June 2023

Superconductivity



The superconductivity was discovered in 1911 by Heike Kamerlingh Onnes at the Cryogenic Laboratory at Leiden University.

At 4.2 K (-296°C), he observed a disappearance of resistivity in mercury. His experiments were made possible by the condensation of helium (1908).

Superconductivity is everywhere



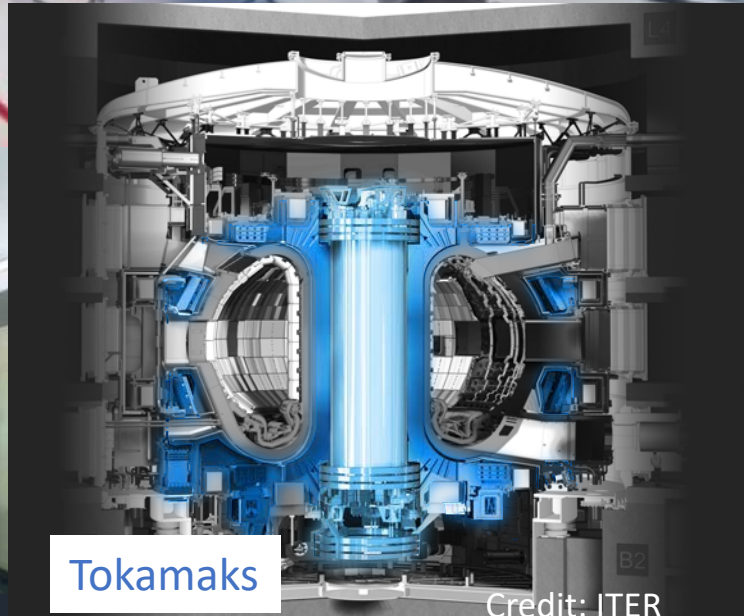
CERN Hadron Collider



MRI scanner



SCMaglev



Tokamaks

Credit: ITER



Quantum Computing

Credit: IBM

Low temperature superconductivity

$$\text{Energy} \propto k_B T$$

$$T \cong 0 \text{ K}$$

Low Motion

High Vision

Massive detectors near the absolute zero

The coldest massive objects in the universe

MiniGRAIL (Leiden)

a spherical gravitational waves detector



1.3 Ton CuAl sphere cooled at 65 mK
 10^{-19} m displacement sensitivity at 3 kHz

A.De Waard et al., Classical Quantum Gravity 21, S465 (2004)
L. Gottardi et al., Phys. Rev. D. 76, 102005 (2007)

CUORE (GranSasso)

Cryogenic Underground Observatory for Rare Events



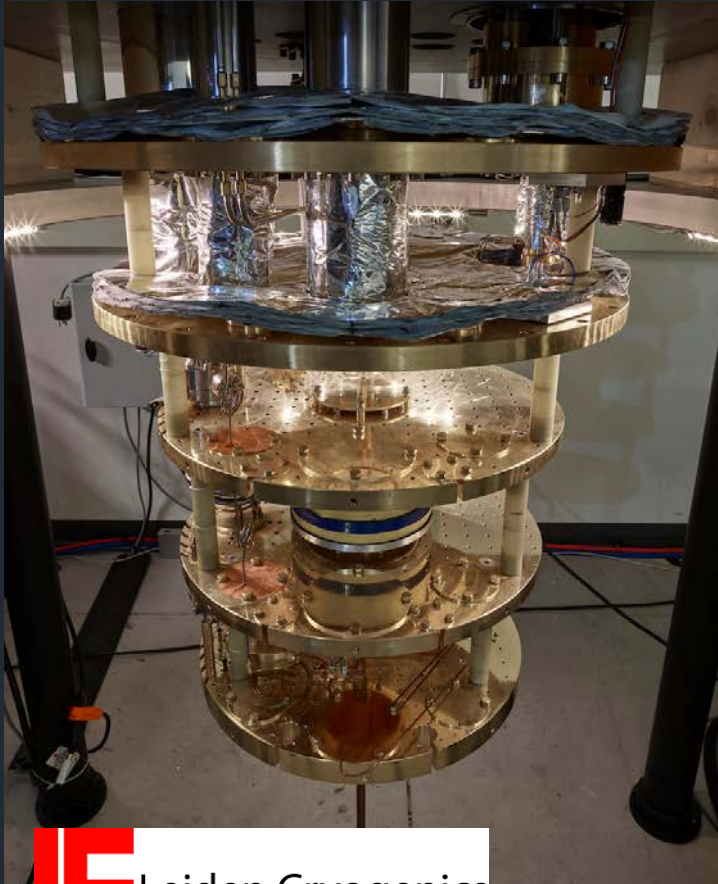
750 Kg TeO₂ crystals + **4 Ton** Cu and lead shields
cooled at 10 mK !

The CUORE Collaboration.. Nature 604, 53–58 (2022)

Ultra Low Temperature refrigeration

The Kamerling Onnes' Legacy

Pioneers and world leaders in ultra low temperature refrigeration



 Leiden Cryogenics
Leader in Low Temperature Techniques



 BLUE FORS

 SRON

Ultra low temperature detectors
in SPACE
is the big challenge



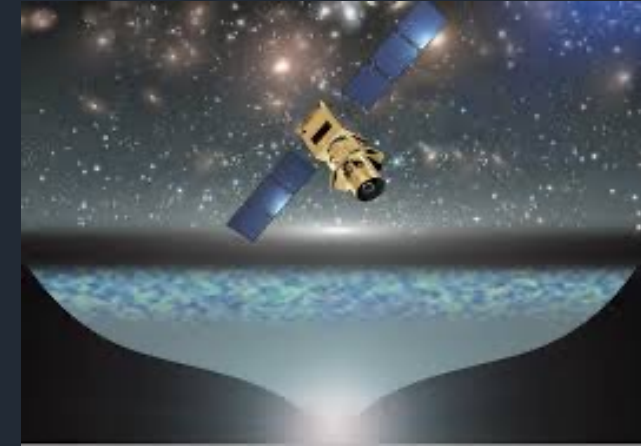
Large array of superconducting detectors for astrophysics in space

Athena/XIFU (~2035)



ESA large mission to explore The Hot and Energetic Universe

LiteBird (~2028)



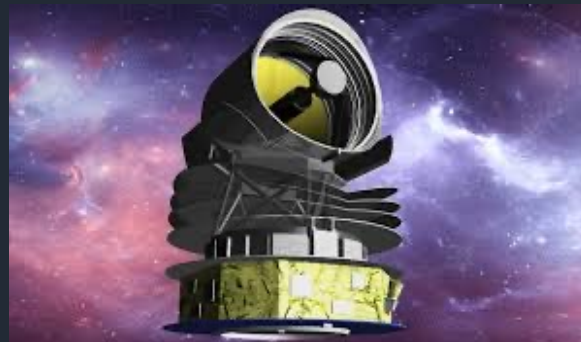
JAXA CMB mission to search for primordial gravitational waves emitted during the cosmic inflation

Hubs



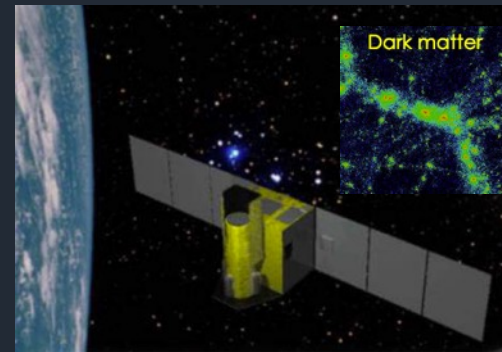
Chinese mission to study 'missing' baryons in the universe

Spica/SAFARI



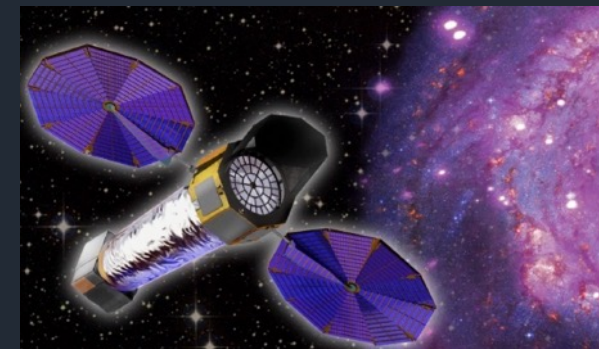
JAXA/ESA infrared space observatories to study formation of stars, planets and galaxies.

SuperDIOS



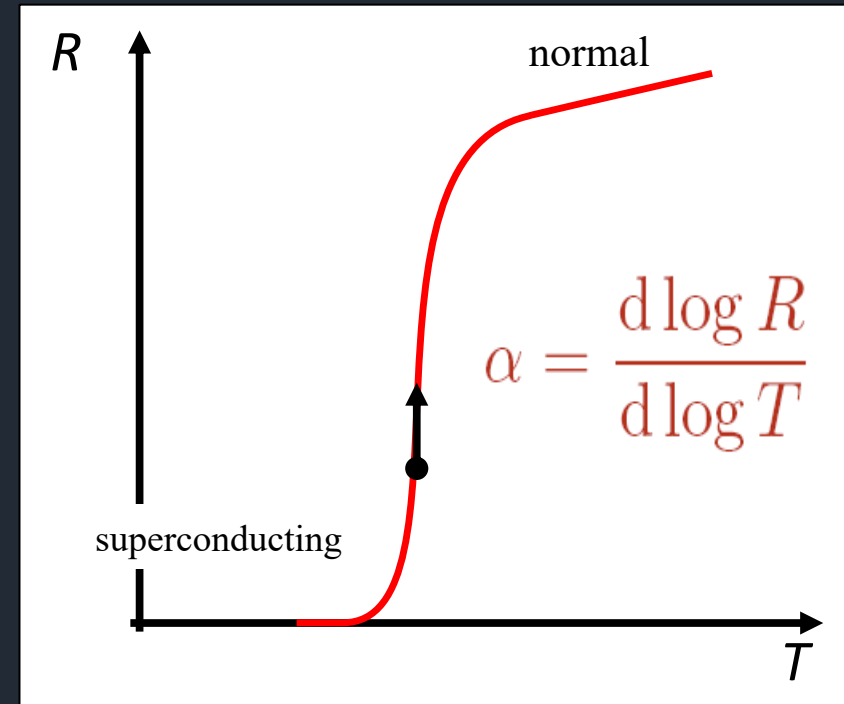
JAXA Diffuse Intergalactic Oxygen Surveyor. An X-ray quantitative exploration of "dark baryon "

Lynx



NASA X-ray Large telescope >100000 pixels

Superconducting Transition Edge Sensors (TESs): very sensitive low-temperature thermometers

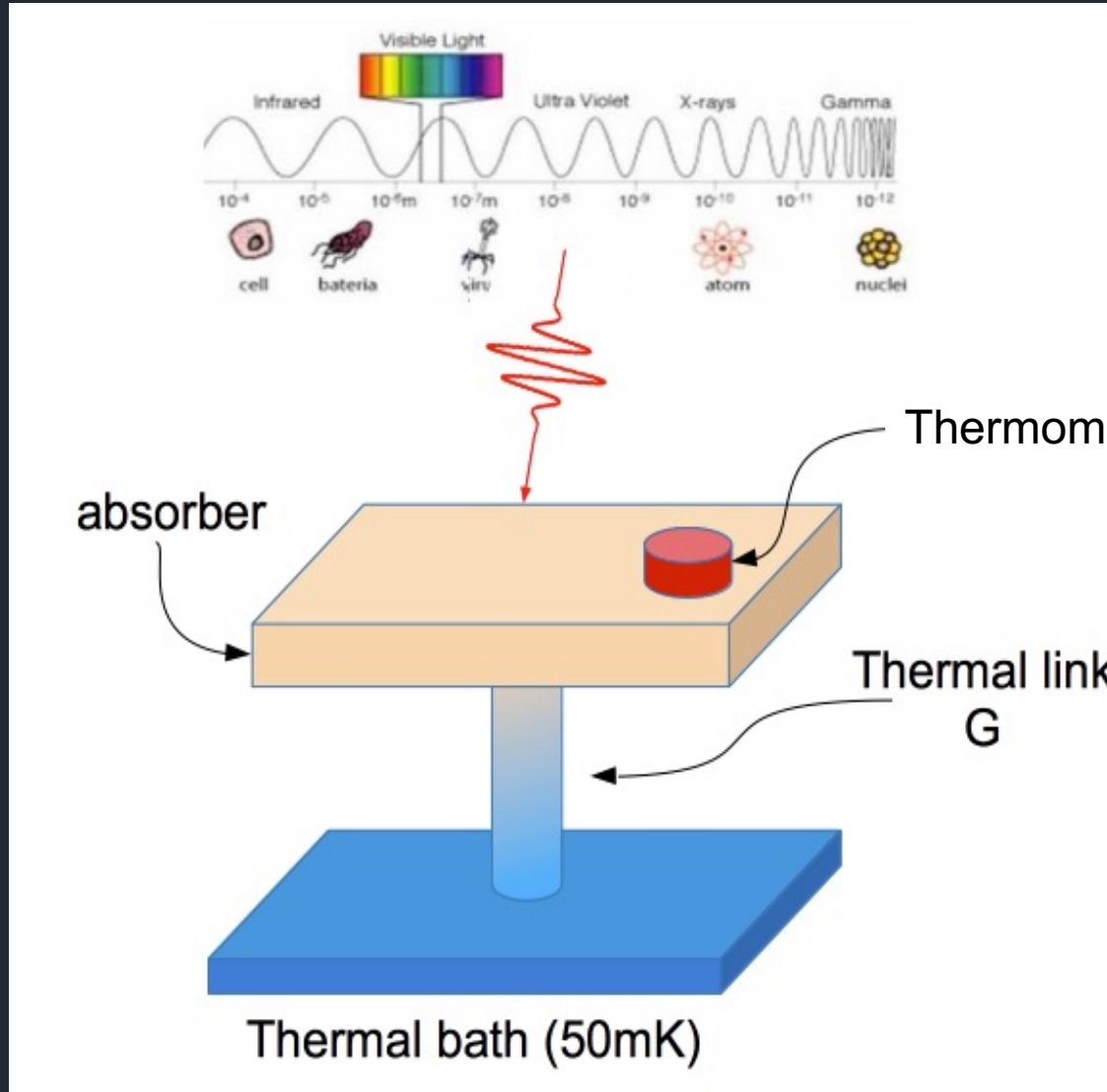


Thin film superconducting bilayer

Ti/Au, Mo/Au, Mo/Cu
 $T_C \sim 90$ mK

D. H. Andrews et al., Phys. Rev. 59, 1045 (1941).
K. D. Irwin, Appl. Phys. Lett. 66, 1945 (1995)

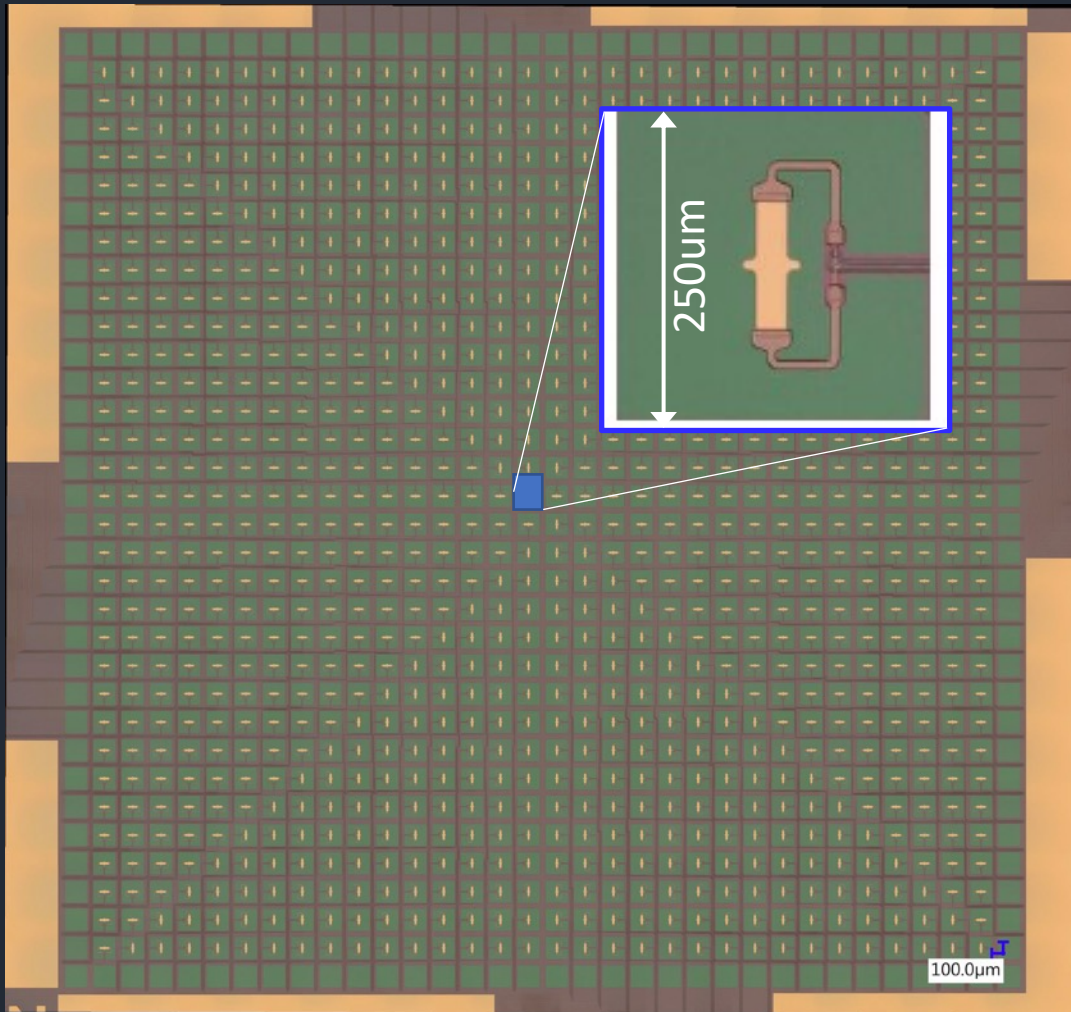
Low temperature micro-calorimeters and bolometers



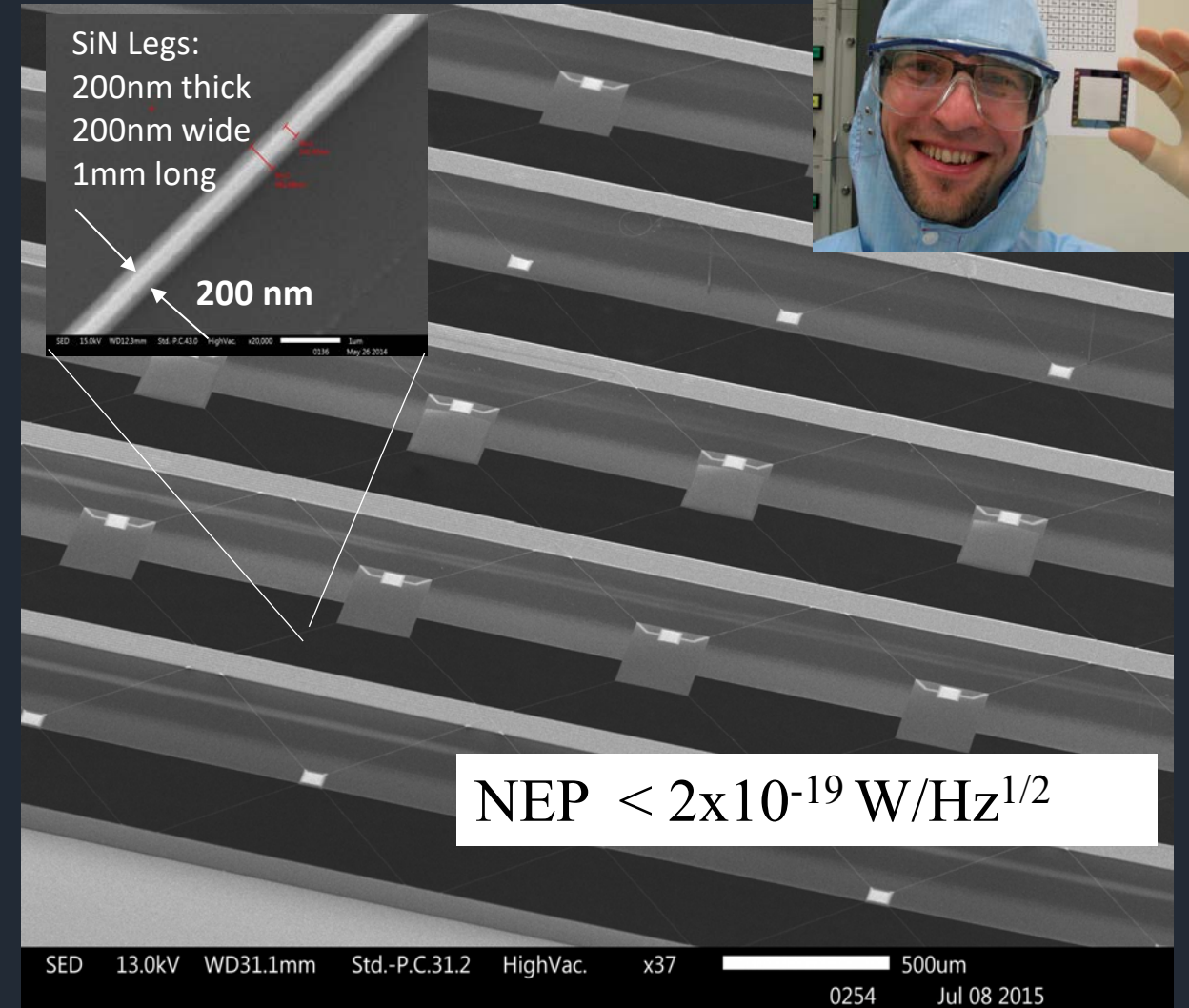
Fundamental elements:
(individually tuneable depending on the applications)

1. Absorber
2. Sensitive thermometer
3. Weak thermal link to bath

X-ray TES microcalorimeters



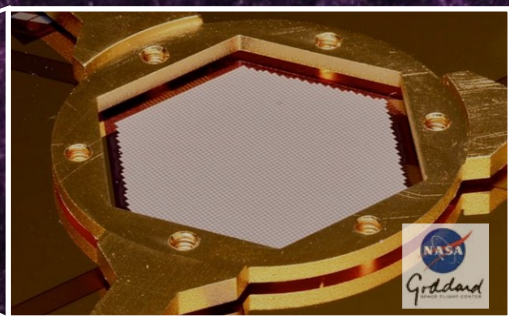
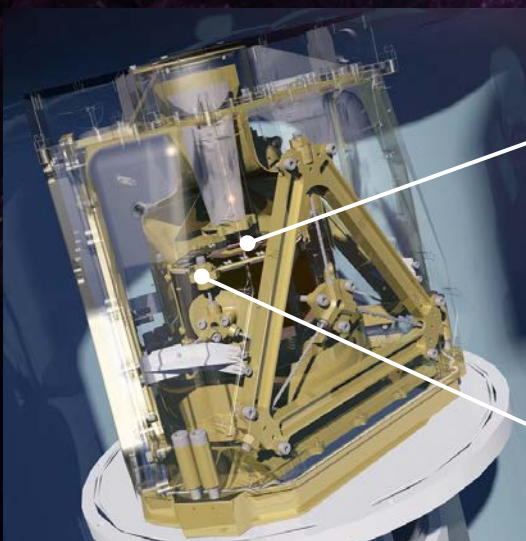
Far-infrared TES bolometers



Future X-ray space observatory



- **ATHENA** is a Large ESA mission to study “*The Hot and Energetic Universe*”, launch in late 2030s.
- The **X-IFU** instruments of the payload is a **cryogenic imaging spectrometer**:
Energy band 0.2 – 12 keV, $dE \sim 2.5$ eV



TES array

**X-ray Integral
Field Unit**

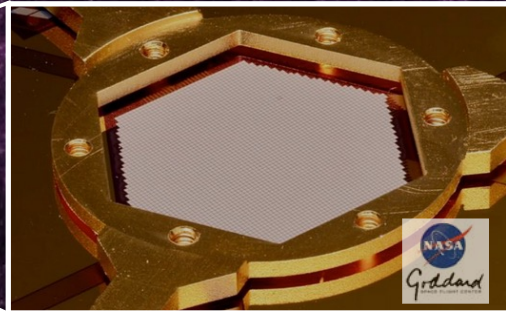
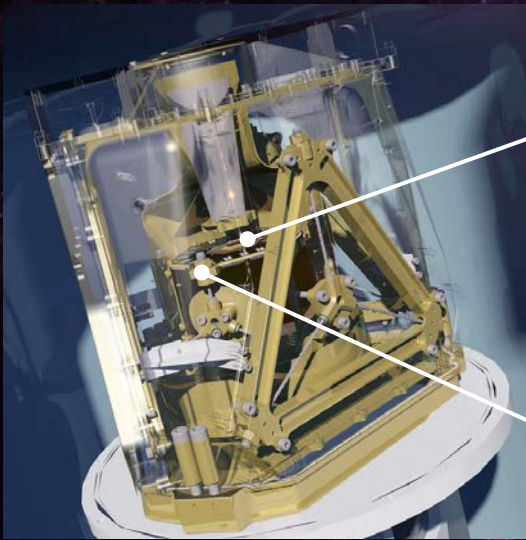


Future X-ray space observatory



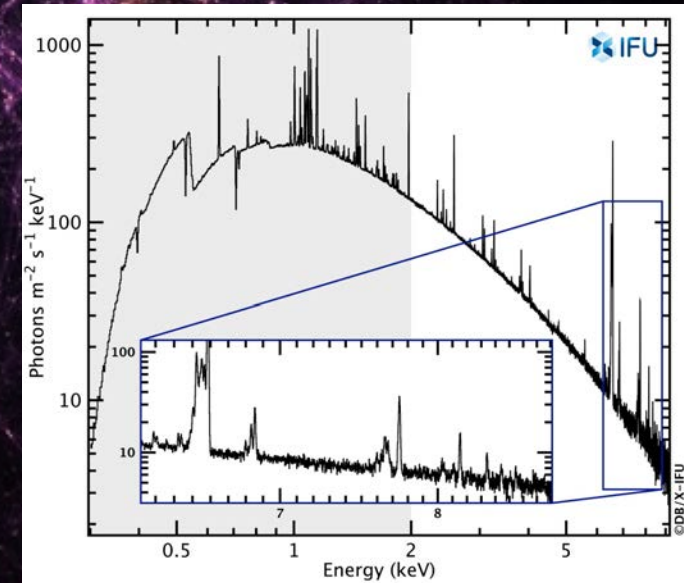
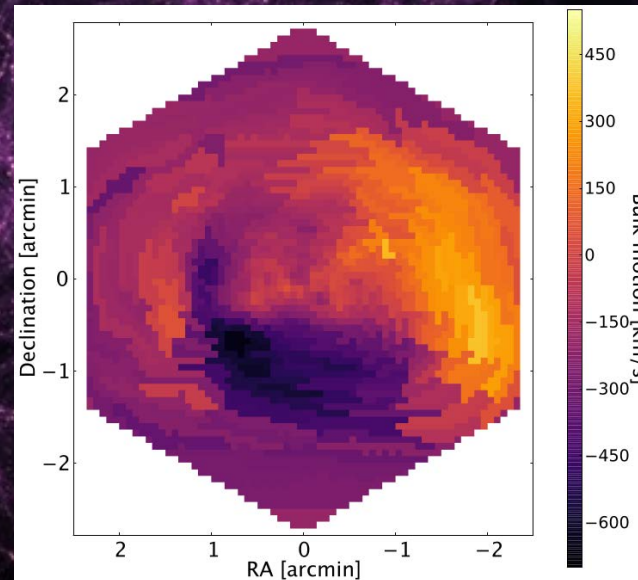
- **ATHENA** is a Large ESA mission to study “*The Hot and Energetic Universe*”, launch in late 2030s.
- The **X-IFU** instruments of the payload is a **cryogenic imaging spectrometer**:
Energy band 0.2 – 12 keV, dE ~ 2.5 eV

The Hot Universe

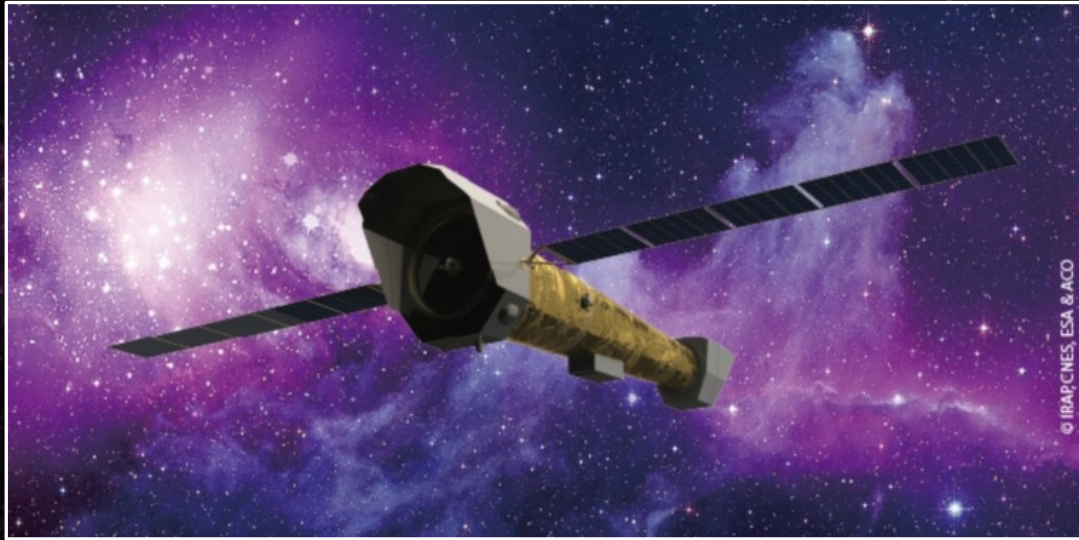


TES array

**X-ray Integral
Field Unit**

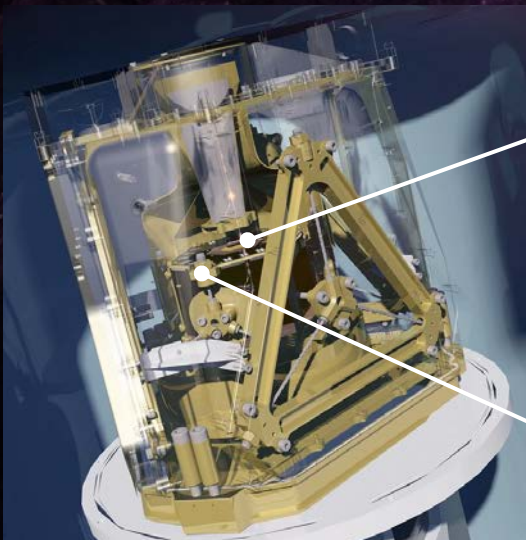


Future X-ray space observatory

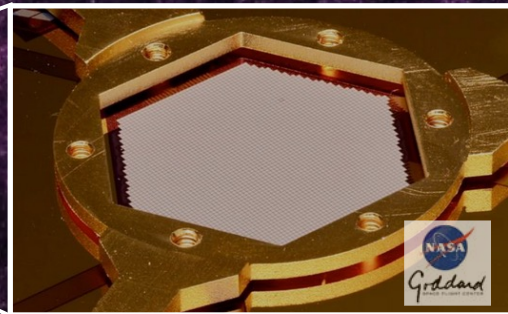


- **ATHENA** is a Large ESA mission to study “*The Hot and Energetic Universe*”, launch in late 2030s.
- The **X-IFU** instruments of the payload is a **cryogenic imaging spectrometer**:
Energy band 0.2 – 12 keV, $dE \sim 2.5$ eV

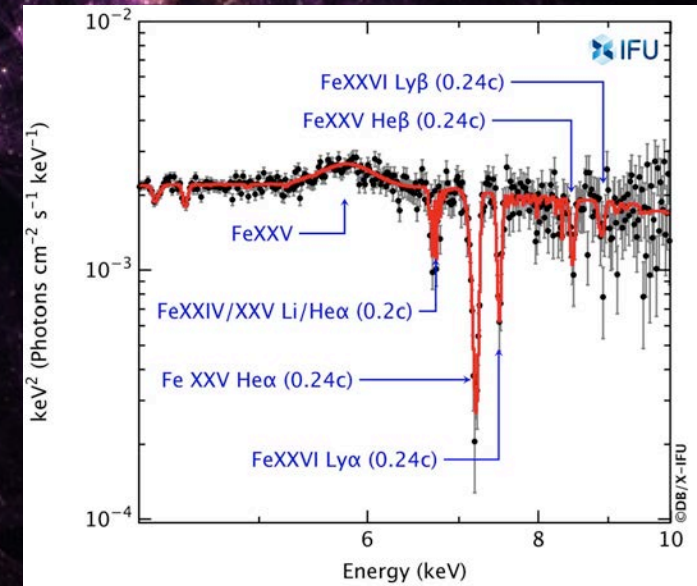
The Energetic Universe



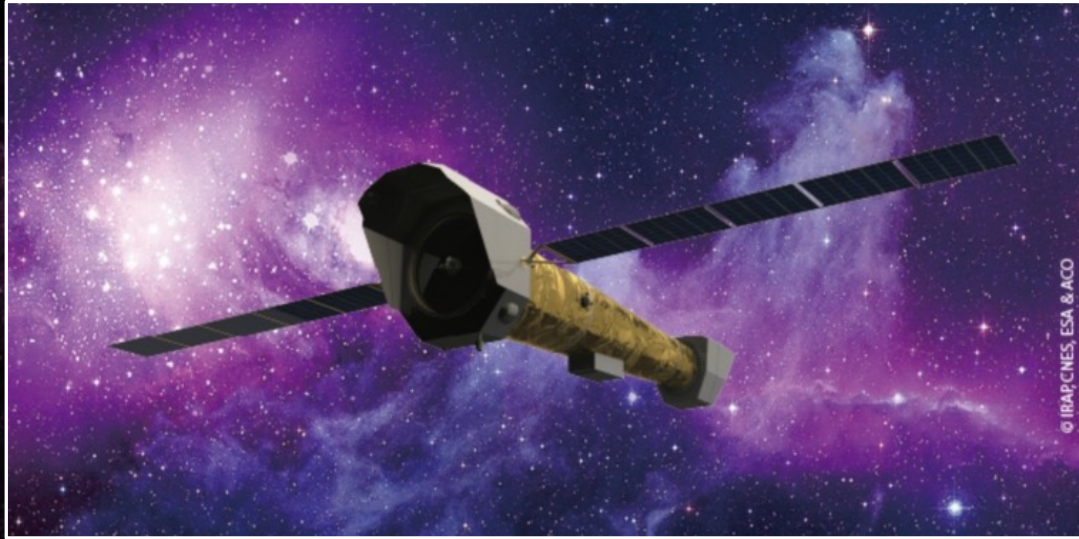
**X-ray Integral
Field Unit**



TES array

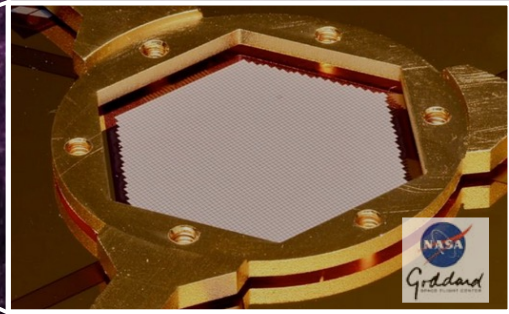
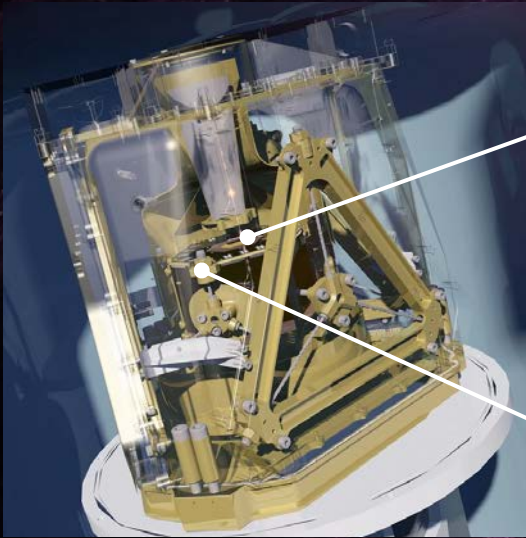


Future X-ray space observatory



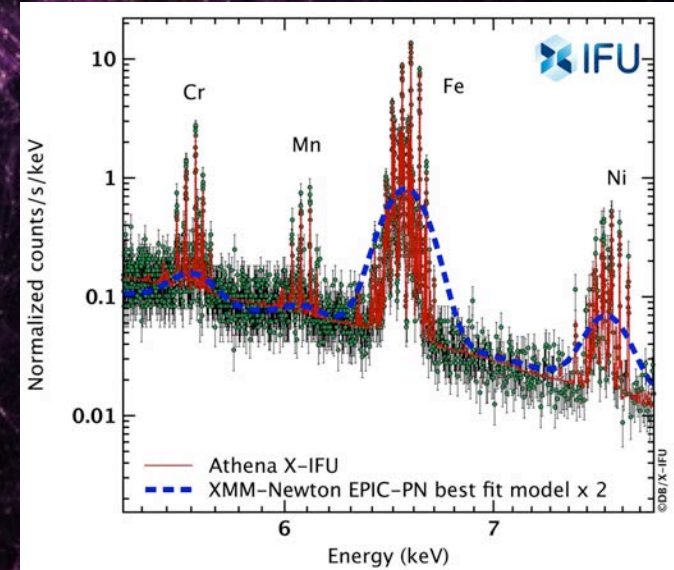
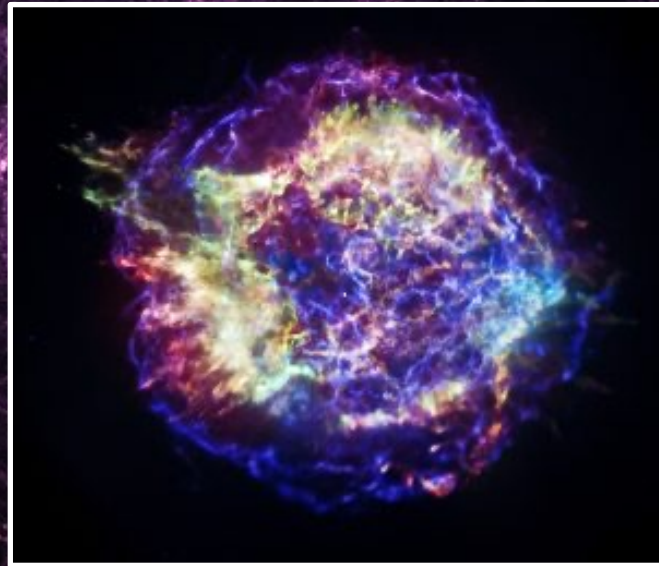
- **ATHENA** is a Large ESA mission to study “*The Hot and Energetic Universe*”, launch in late 2030s.
- The **X-IFU** instruments of the payload is a **cryogenic imaging spectrometer**:
Energy band 0.2 – 12 keV, $dE \sim 2.5$ eV

Observatory Science

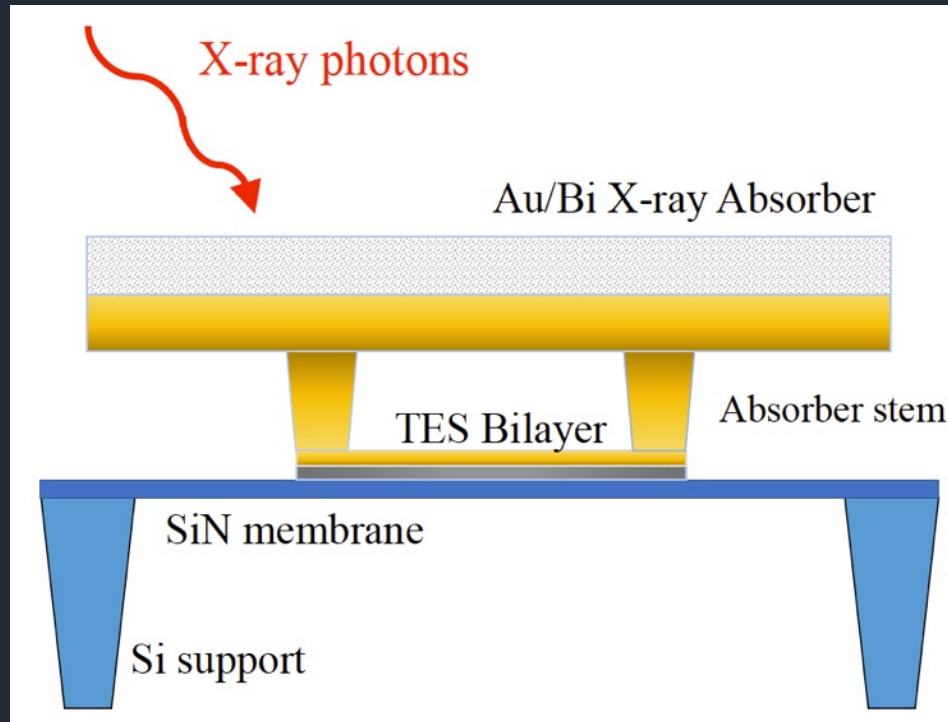


TES array

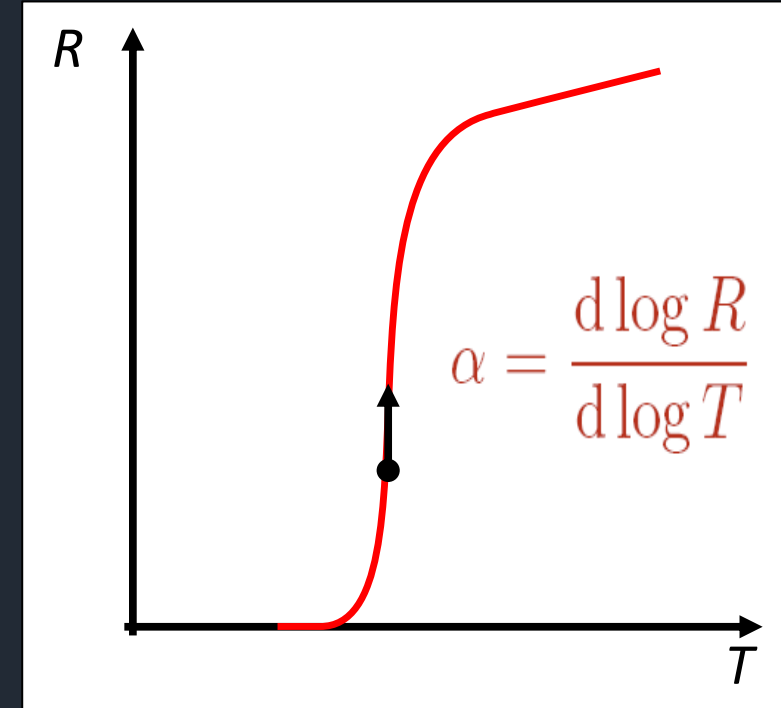
X-ray Integral
Field Unit



Superconducting Transition Edge Sensors



- Low temperature detectors $T_c \sim 90$ mK
- Sharp transition $\alpha \sim 500-1000$
- Small absorber (low heat capacity C)
- Limited dynamic range $E_{lin} \sim C/\alpha$

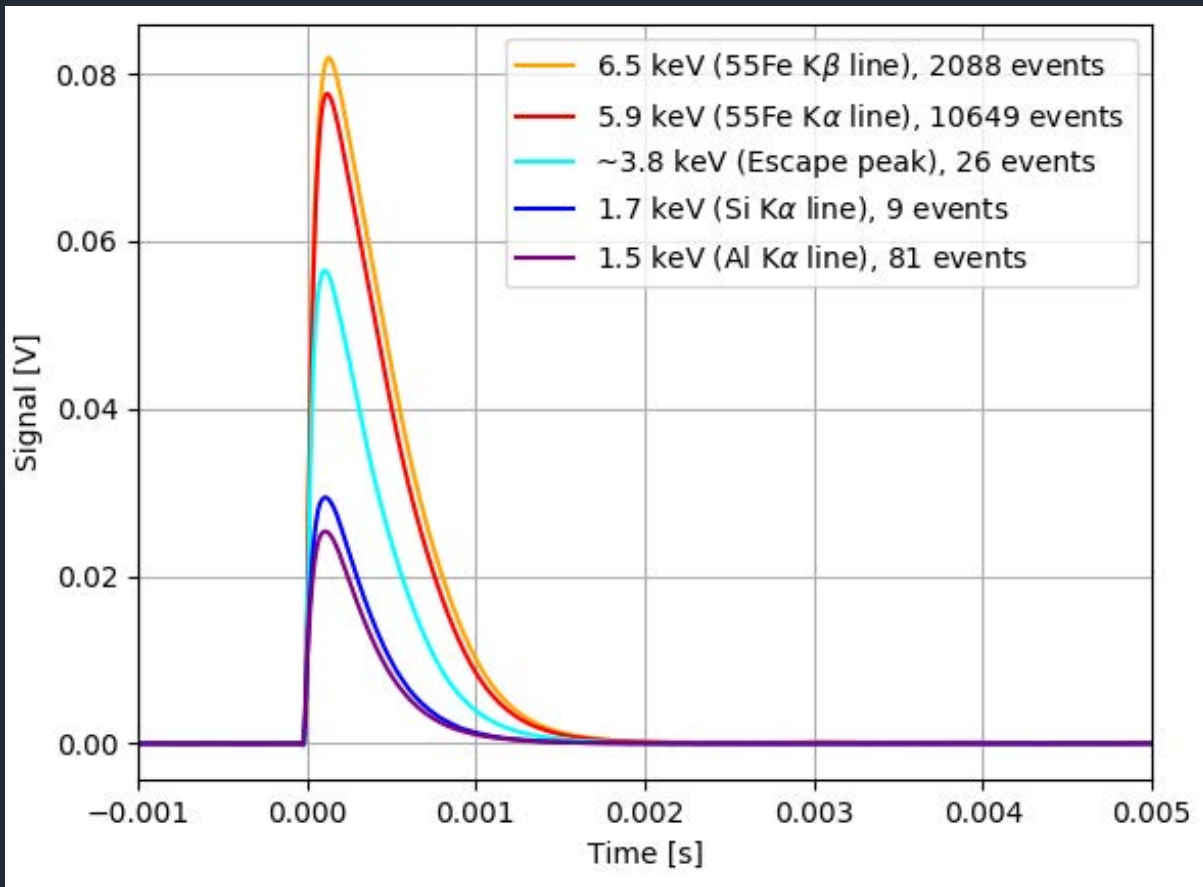


Energy resolution:

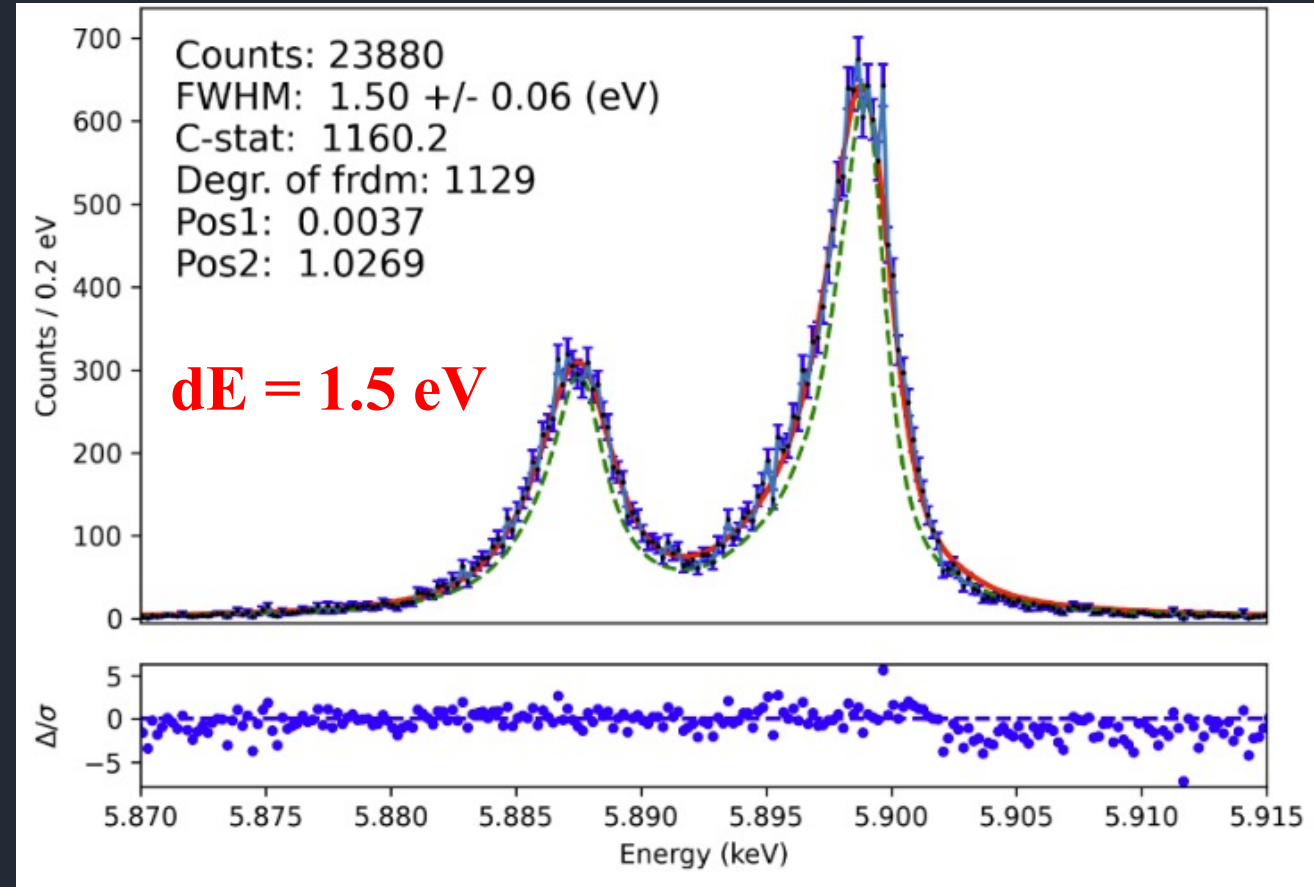
$$E_{FWHM} \sim 2.355 \sqrt{\frac{4k_B T_c^2 C}{\alpha}}$$

K.Irwin and G. Hilton In Cryogenic Particle Detection; Enss,C. Springer, 2006
J. Ullom and D. Bennett, Superc.Sci.Technol. 28, 084003, 2015
L. Gottardi and K. Nagayashi, Applied Sciences 11 (9),3793, 2021

TES X-ray microcalorimeters



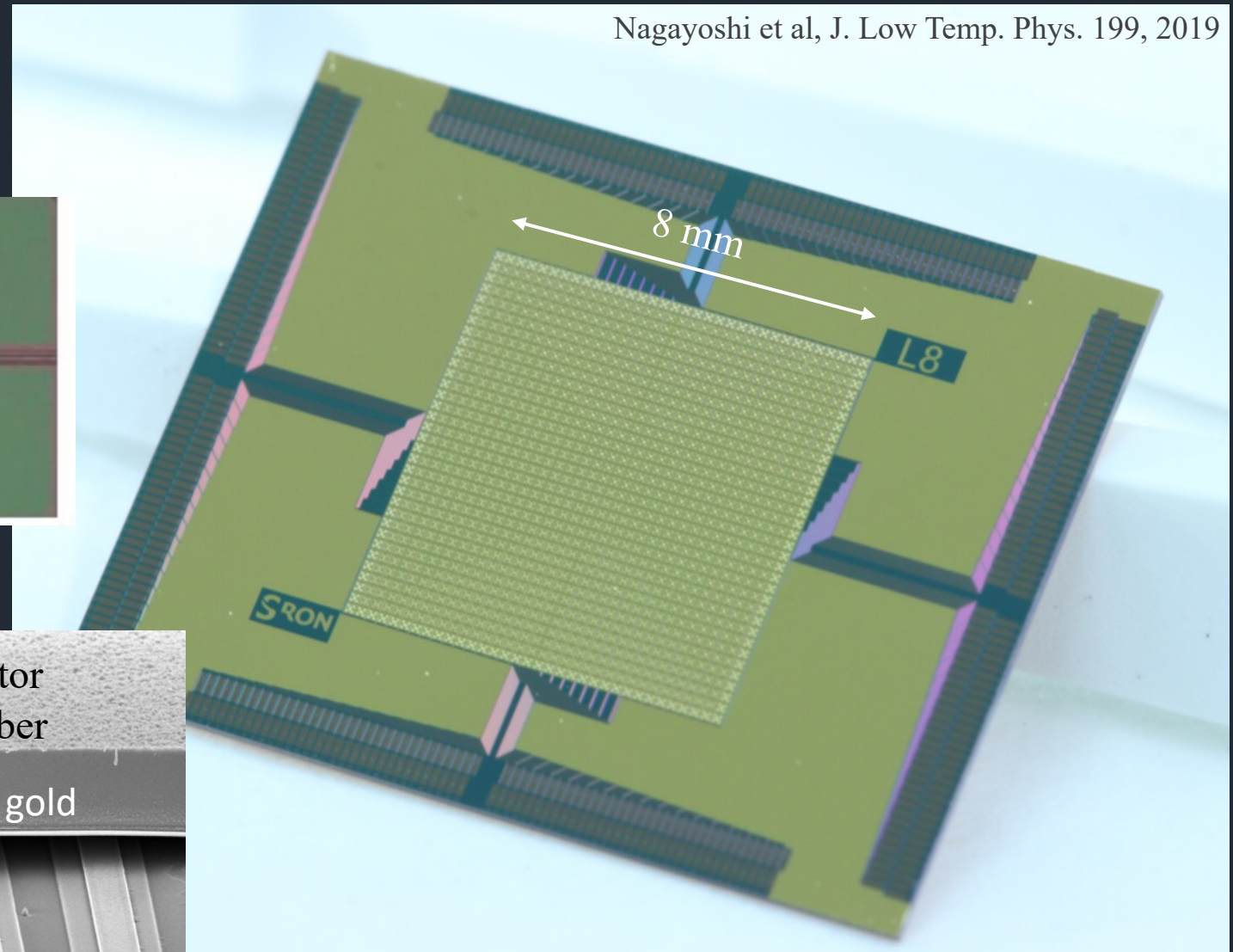
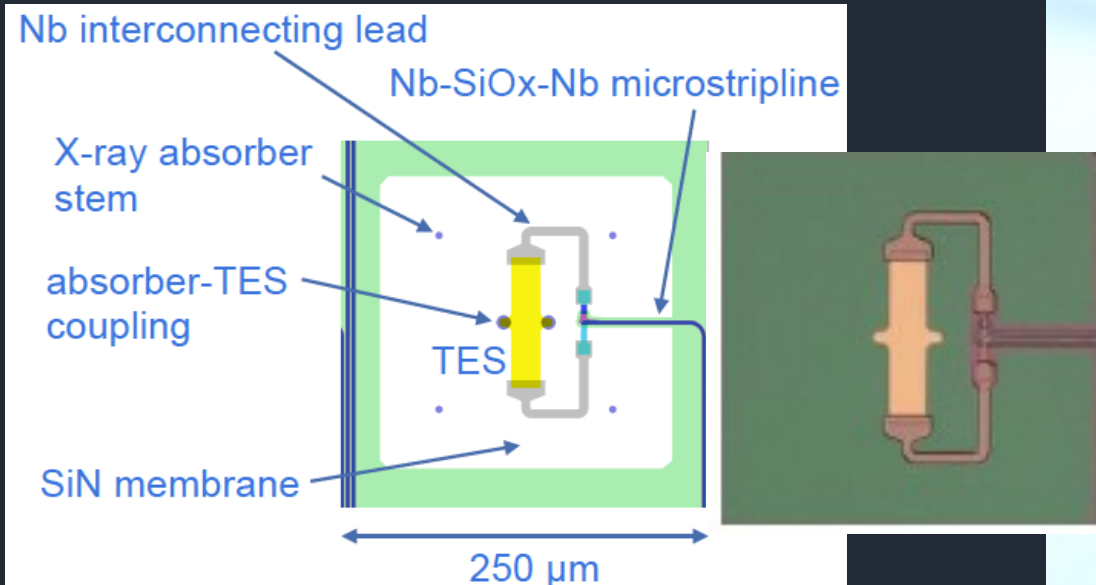
Single photon detectors



High resolving power

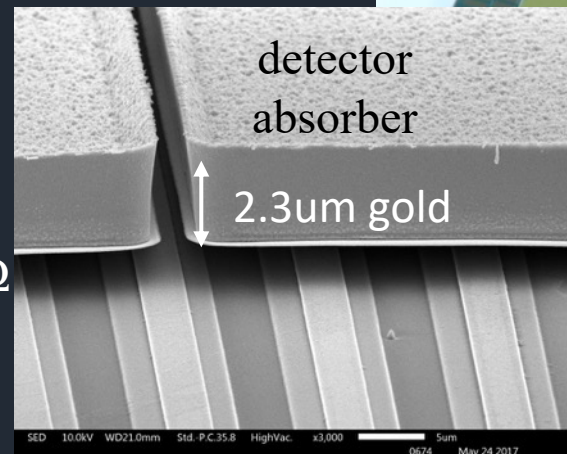
X-ray TES microcalorimeters for XIFU

Nagayoshi et al, J. Low Temp. Phys. 199, 2019



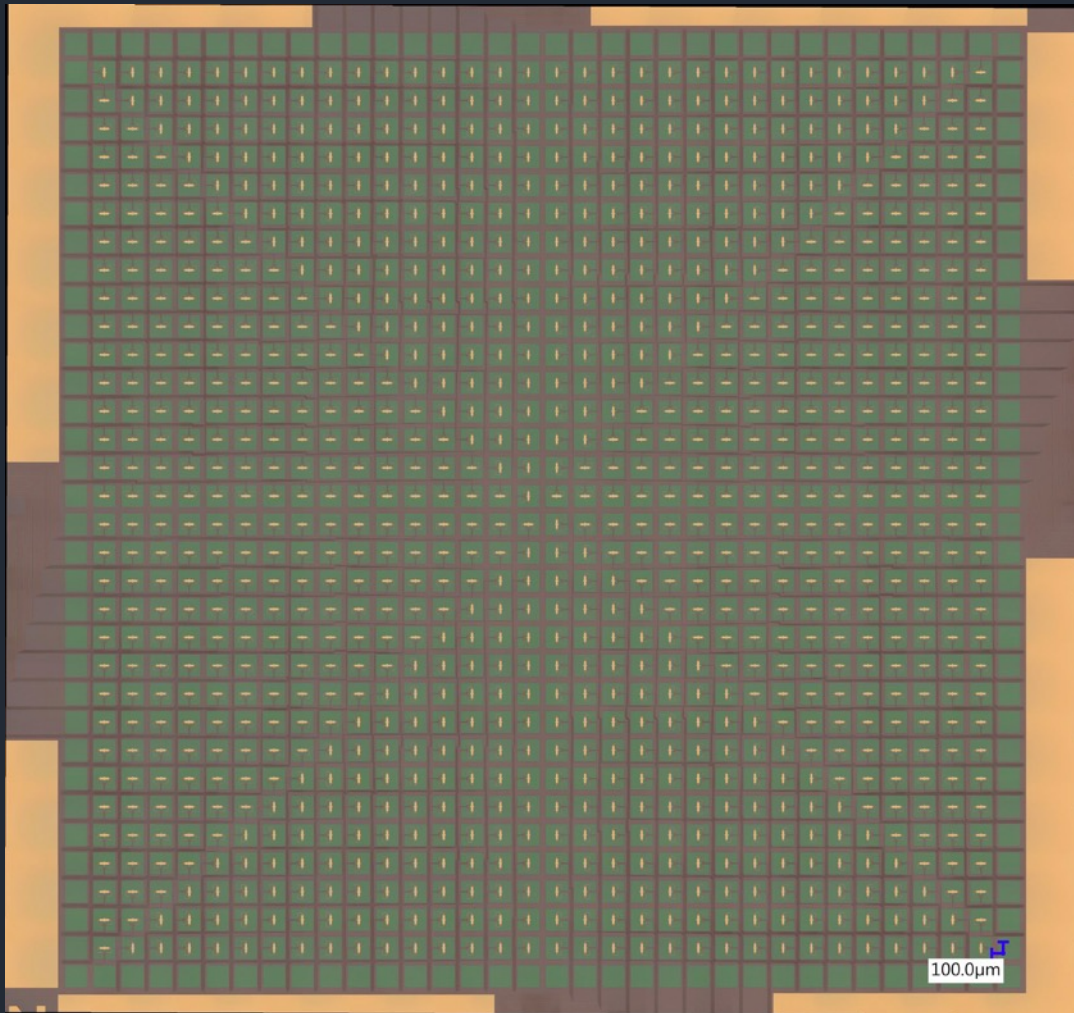
SRON TES:

- Ti(35nm)Au(200nm) bilayer
- $T_c \sim 90\text{mK}$
- Normal resistance $\sim 100\text{-}300\text{m}\Omega$

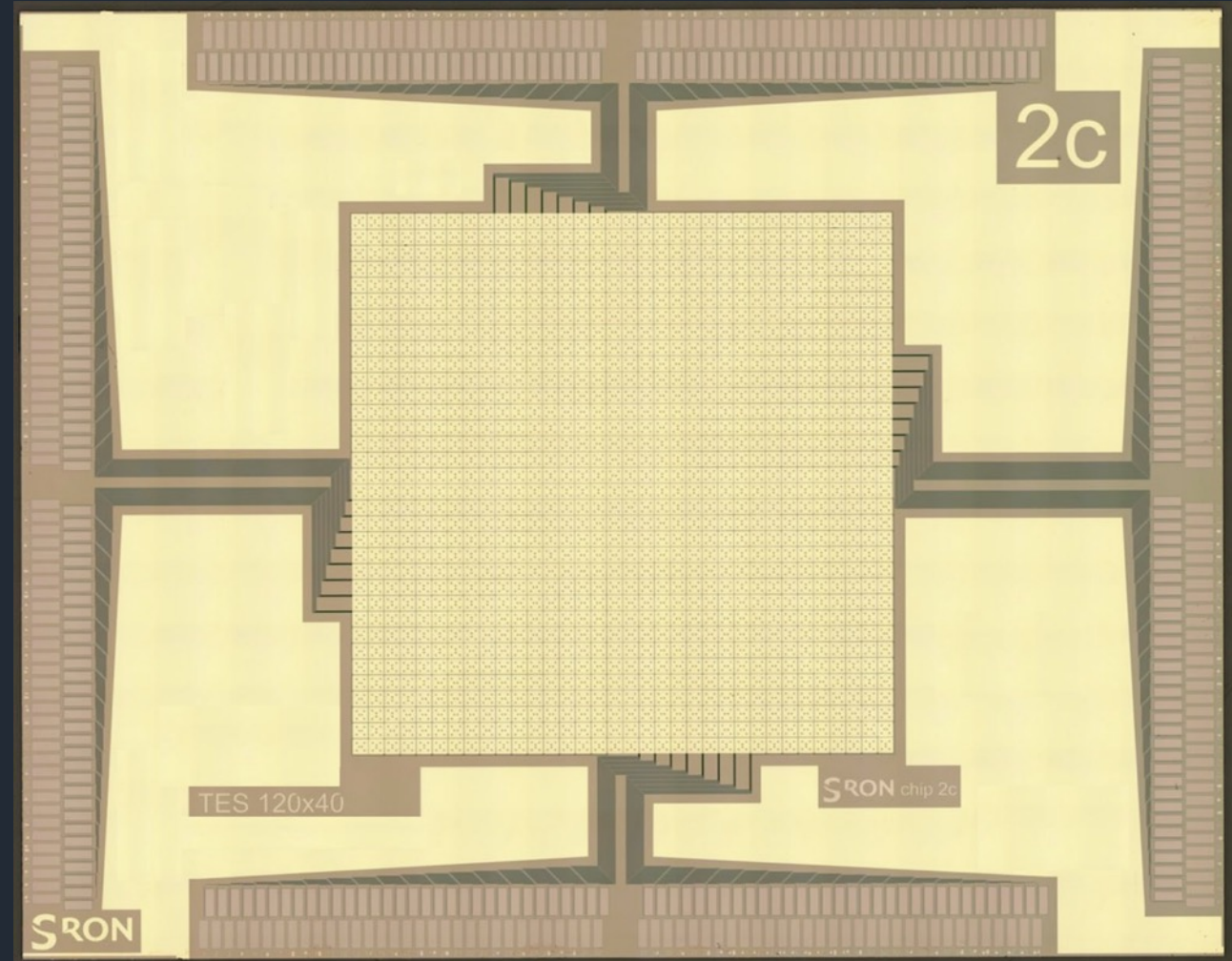


SRON 32x32 TES pixels array with Au absorbers

X-ray TES microcalorimeters for XIFU

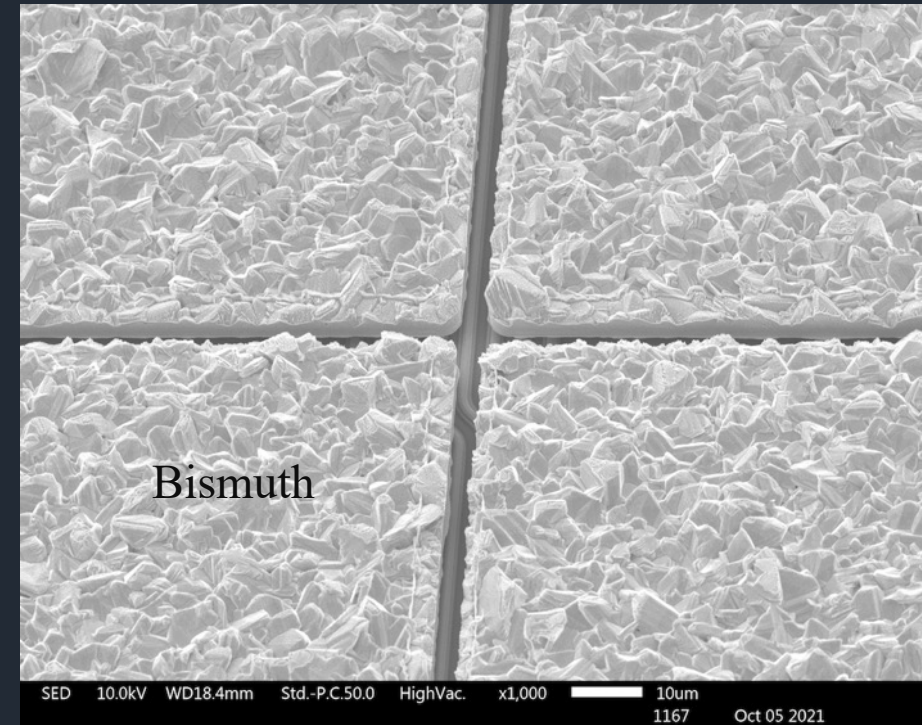
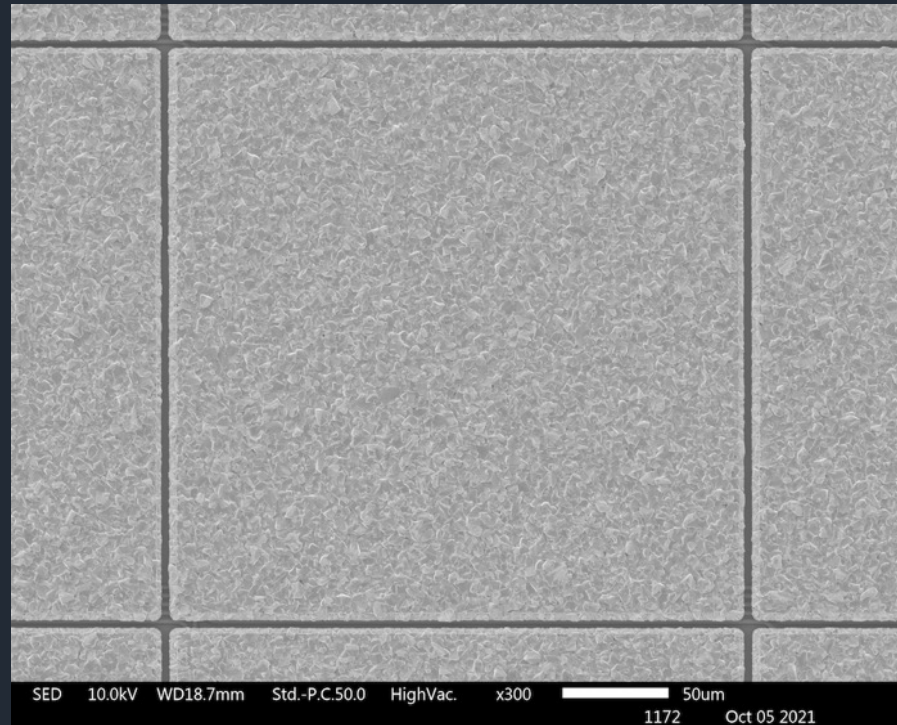
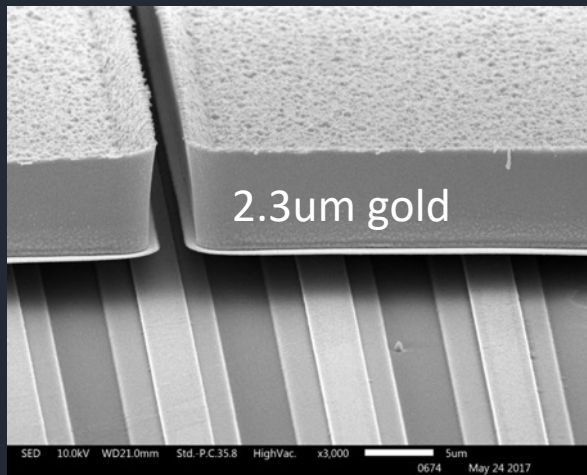
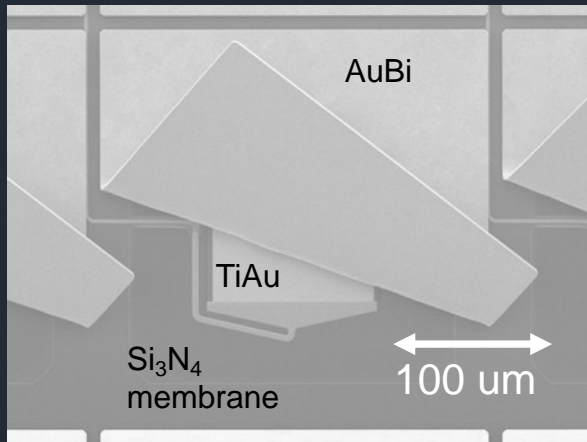


Micrograph SRON 32x32 array before absorber deposition



Final chip SRON 32x32 X-ray microcalorimeters

High filling factor and high quantum efficiency using Au/Bi

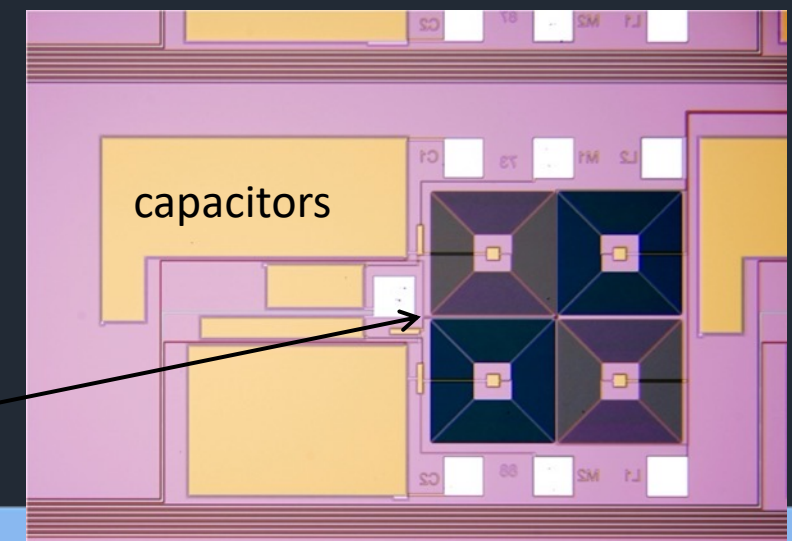
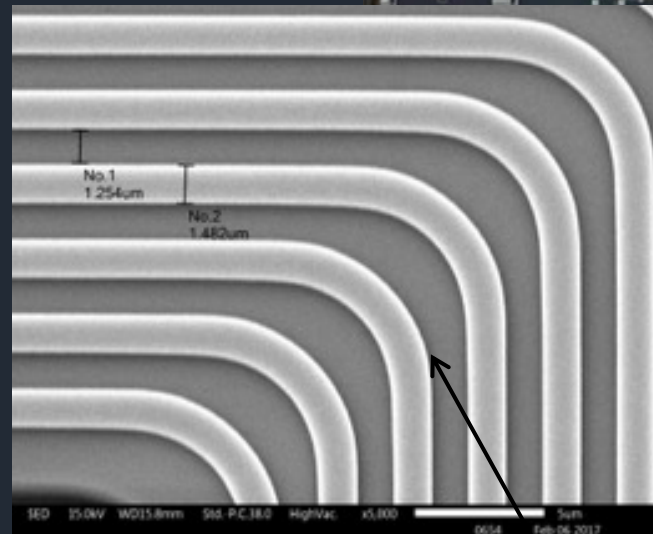
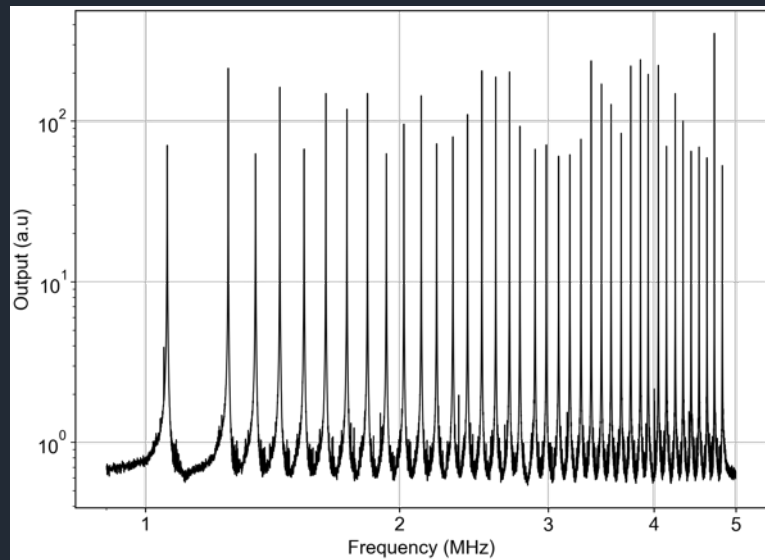


- 2.3 μm electroplated Au absorber QE ~ 83%
- Adding 4 μm electroplated Bi QE → ~ 97%

Superconducting high- Q MHz LC filters

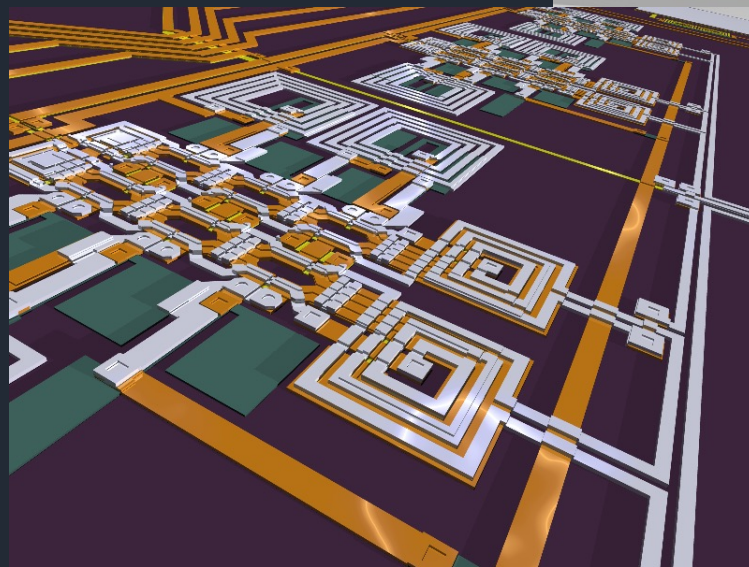
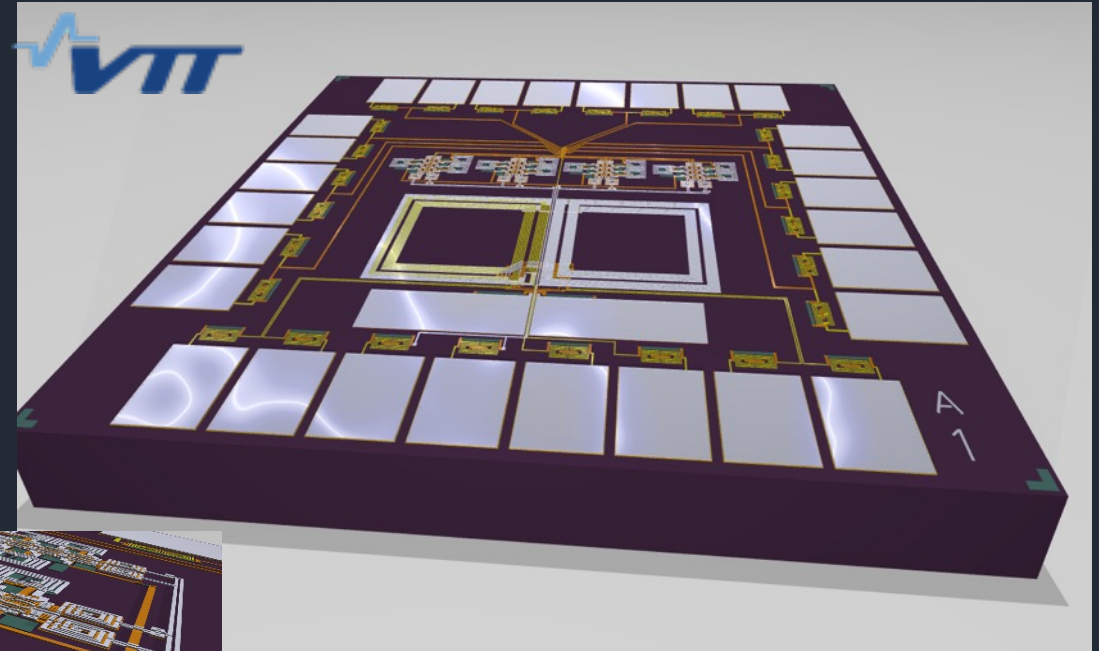
M. P. Bruijn, et al., J. of Low Temp. Phys. 167, 695 (2012).

- Thin film Nb superconducting technology
- Coplanar wiring
- Low loss **amorphous silicon** capacitors
- Gradiometric design to minimize pixel crosstalk
- **High yield (> 97%)**

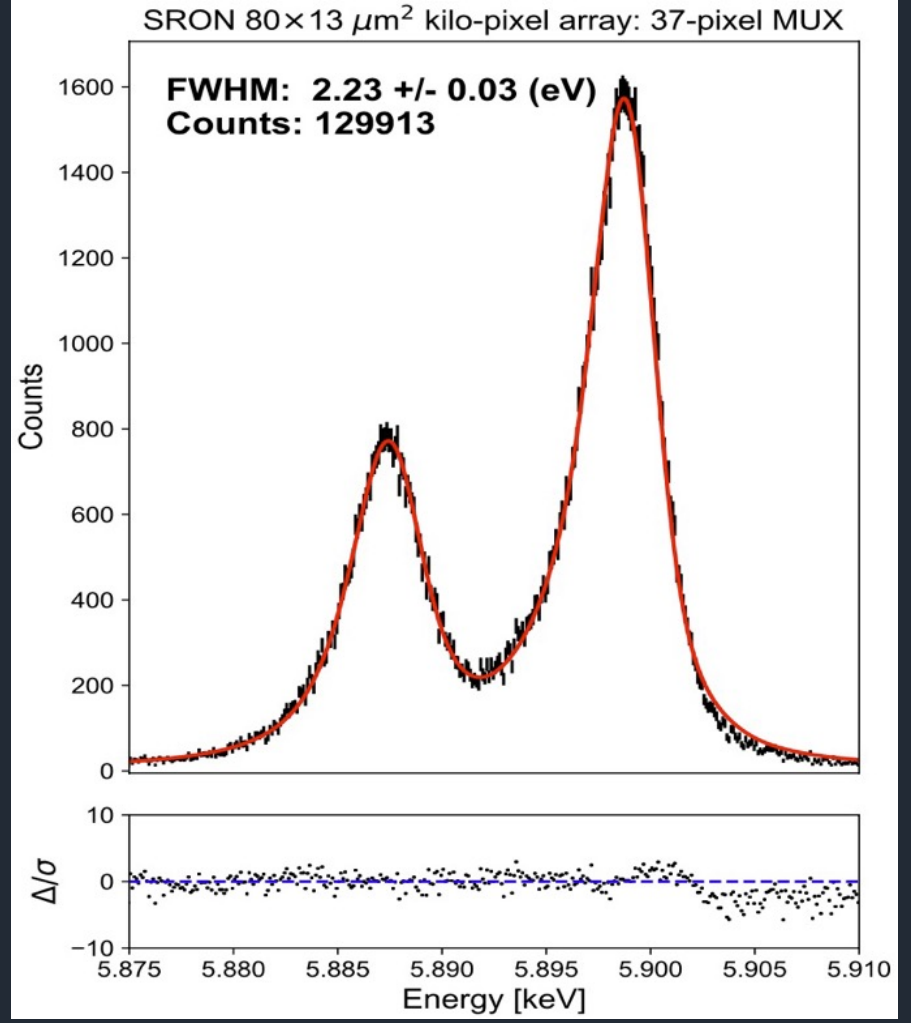
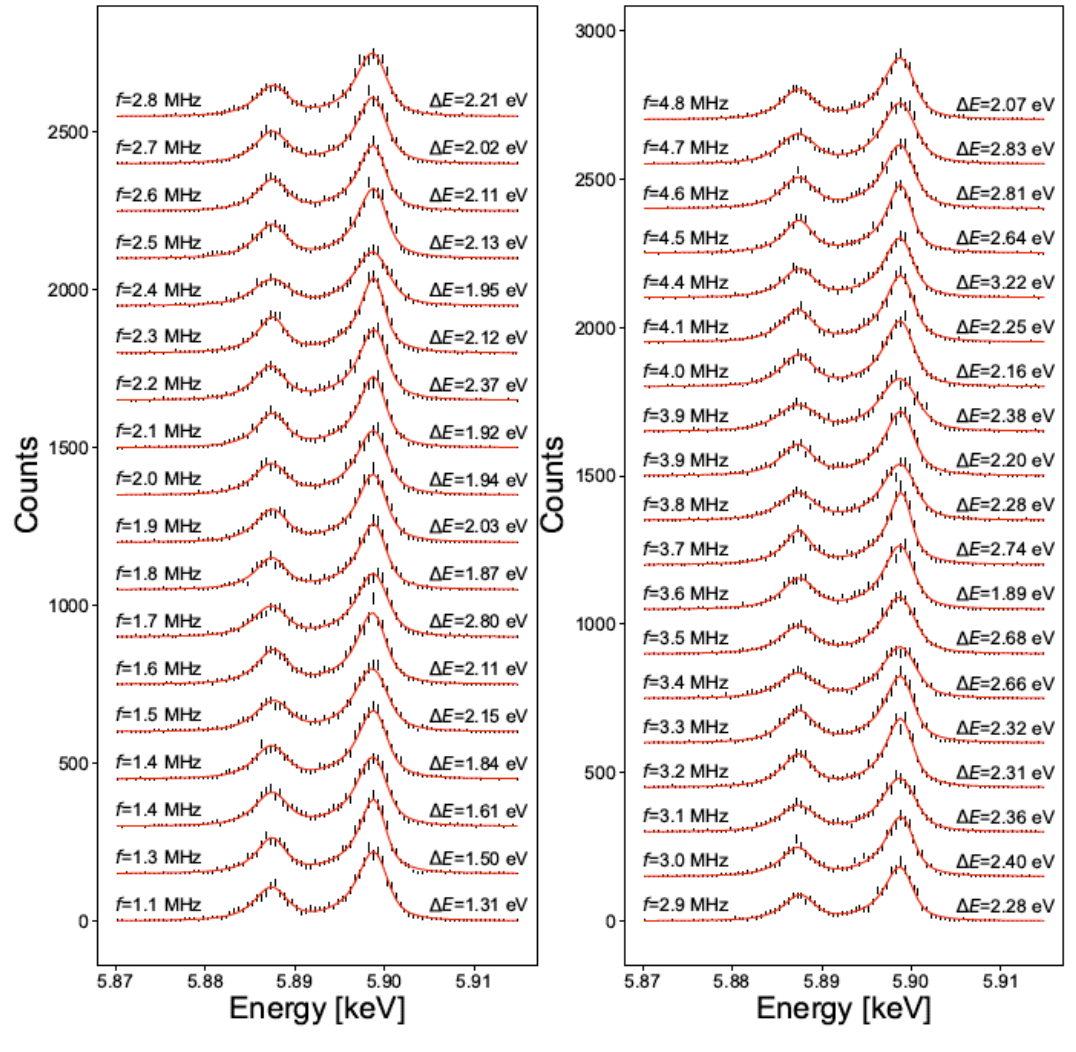


Superconducting QUantum Interference Devices

- Nearly Quantum Limited current amplifiers
- Low power dissipation ($<1\text{nW}$) to be used in space
- Gradiometric design
- Low input inductance to minimize signal cross talks



Frequency division multiplexing demonstration



- Energy resolution fulfills XIFU requirements
- SRON technology is ready for scaling up towards a real instrument



Akamatsu et al. JLT 199, 737 (2020)
Vaccaro, HA et al. RSI 92, 033103 (2021)
van der Hulst, HA et al. RSI 92, 7 (2021)
H. Akamatsu et al. APL 119,18, (2021)

SRON Leiden Cleanroom



E-beam evaporator, cluster tool, LLS



Lithography area

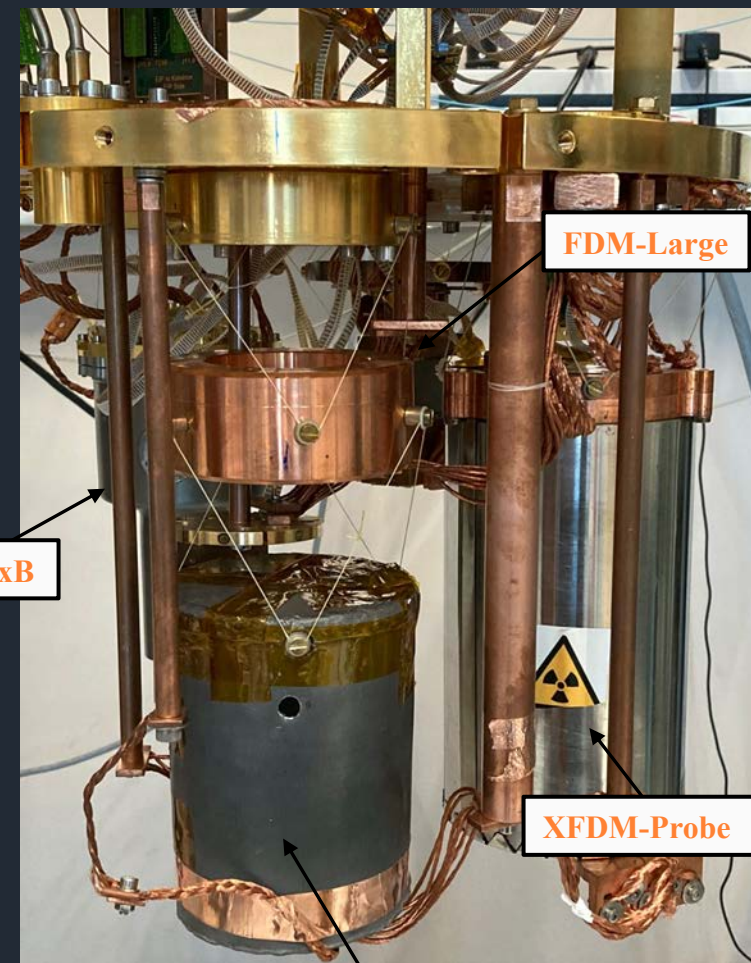
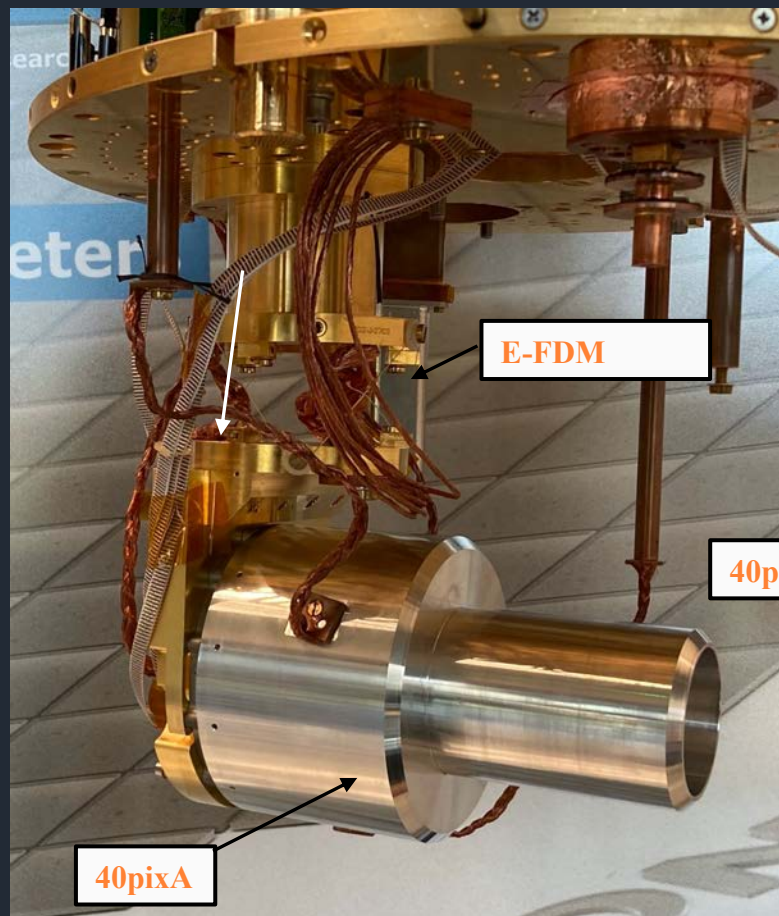


Politeknik: cluster sputter tool



Evatel sputter tool: Nb,Ta,Cu,Al,SiO₂,NbTiN,TiN

Overview of SRON Leiden lab

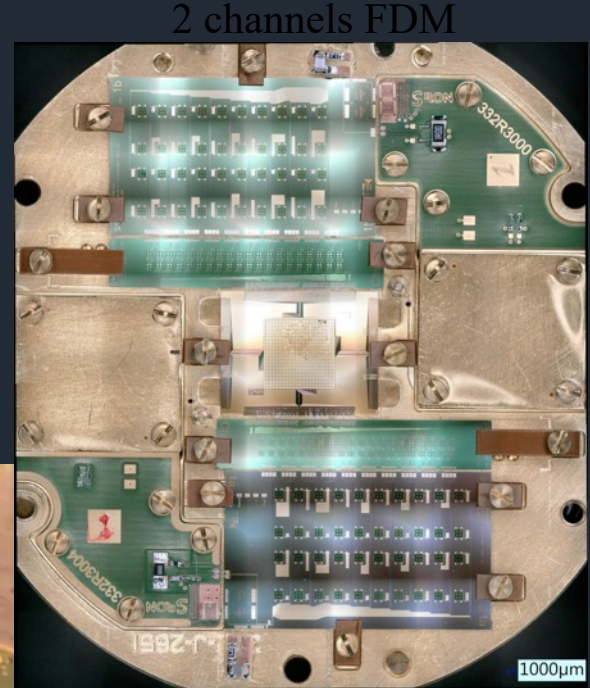
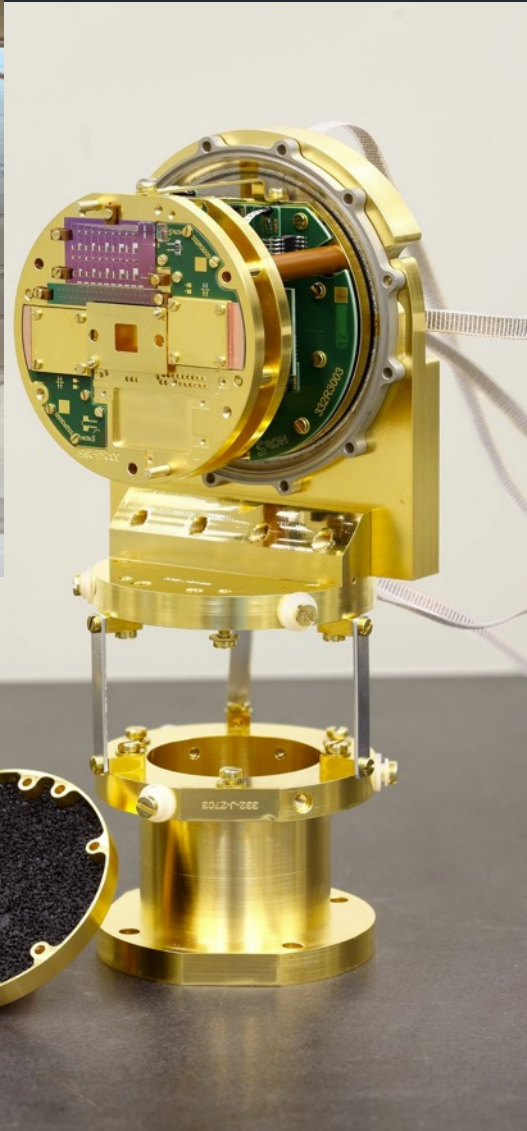


2 Leiden Cryogenics DR with multiple FDM set-ups

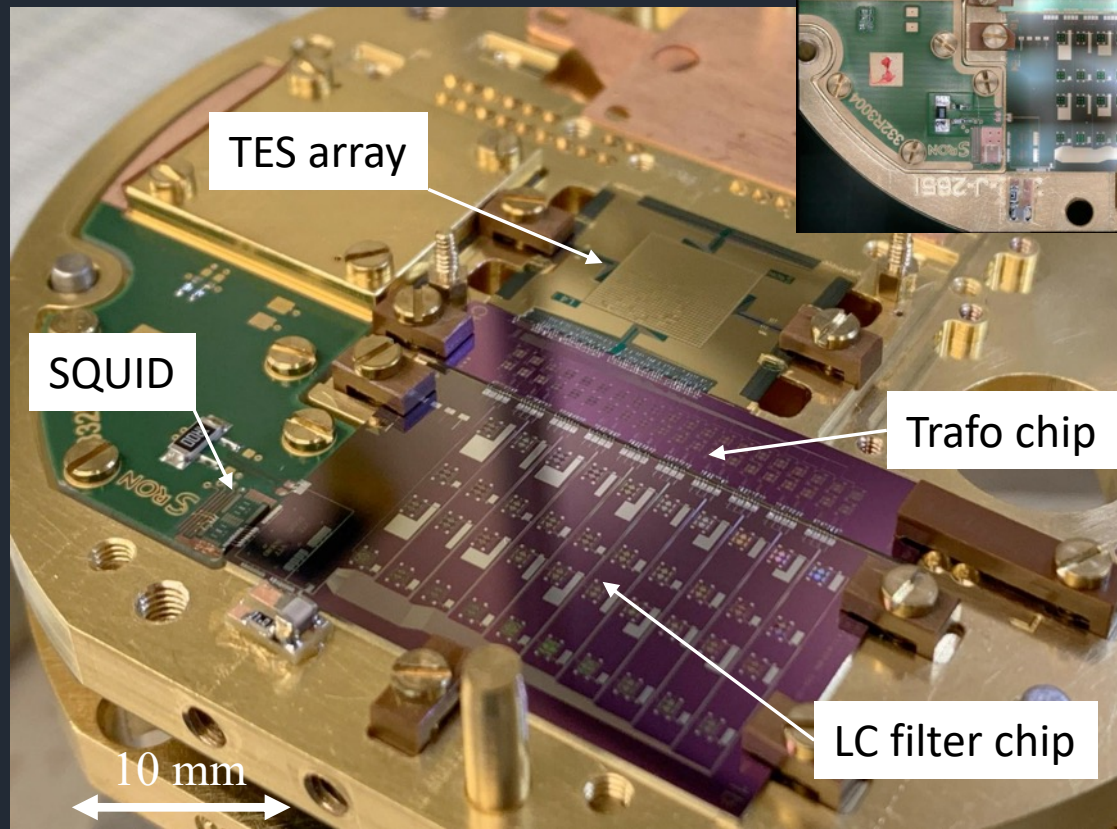
40 pixels FDM demonstrator



Superconducting shield



2 channels FDM



TES array

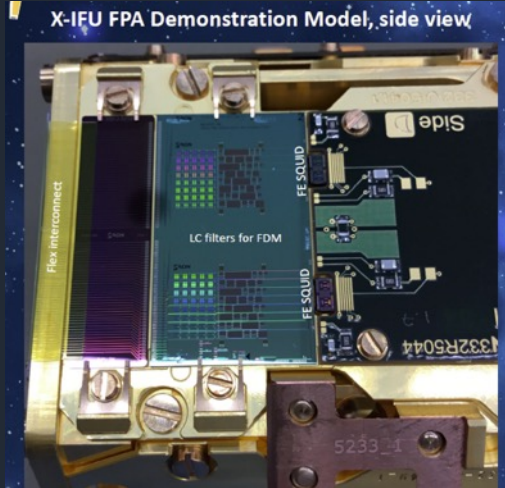
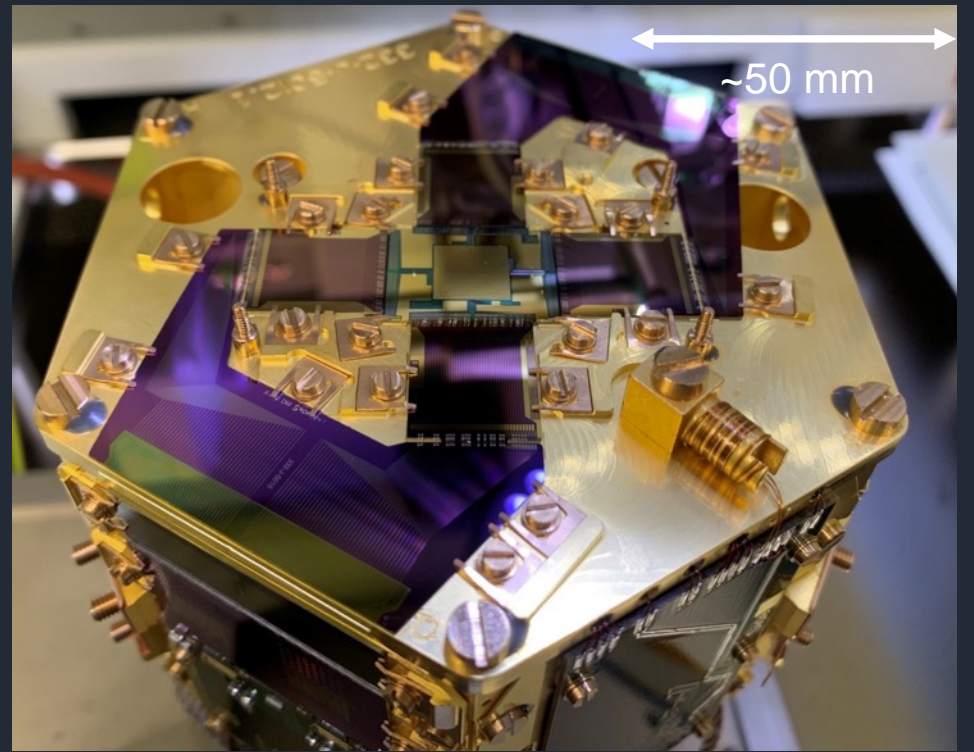
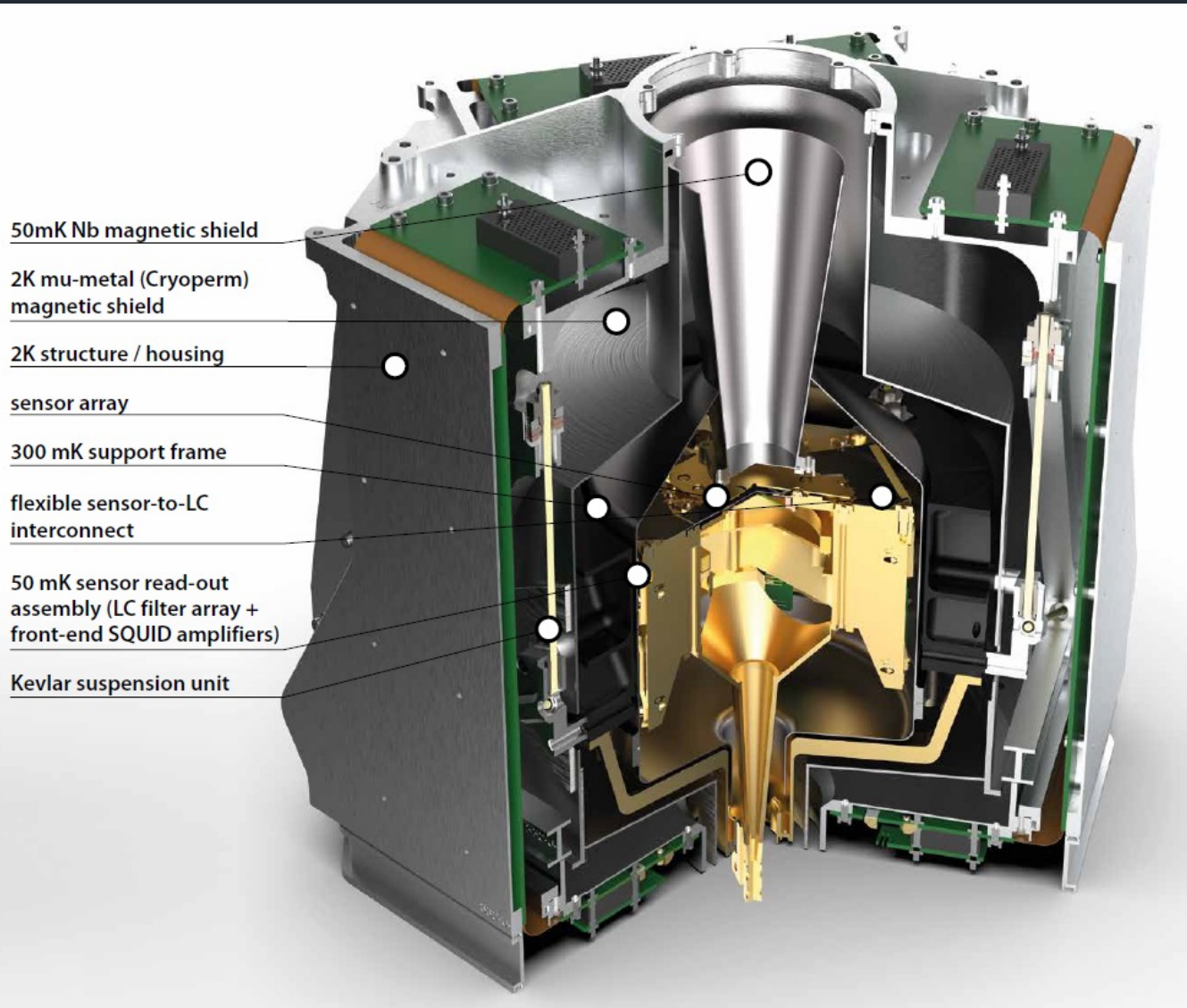
SQUID

Trafo chip

LC filter chip

10 mm

FPA-DM for Athena/XIFU

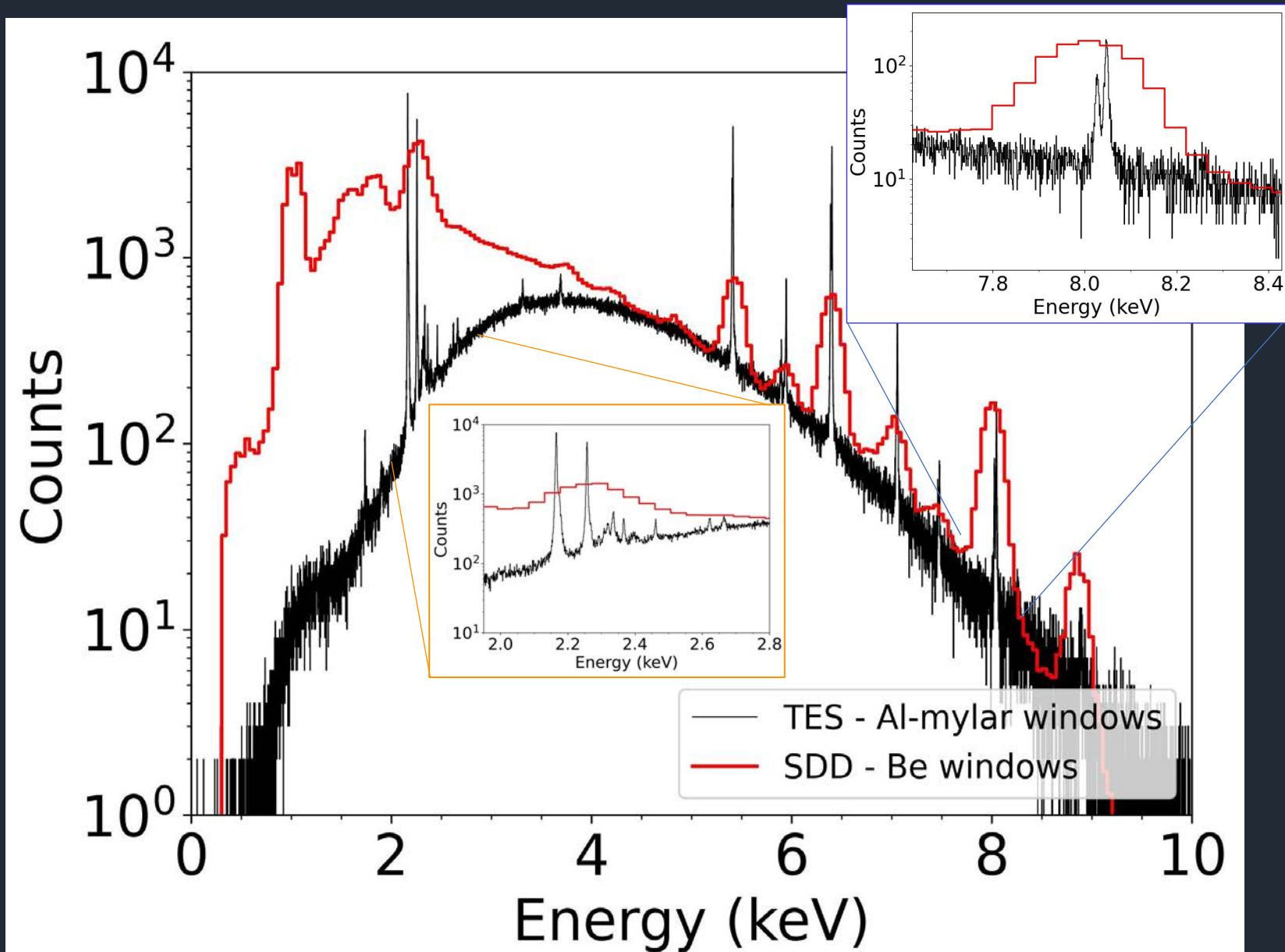


TES based high-resolution X-ray spectrometry
for ground-based applications

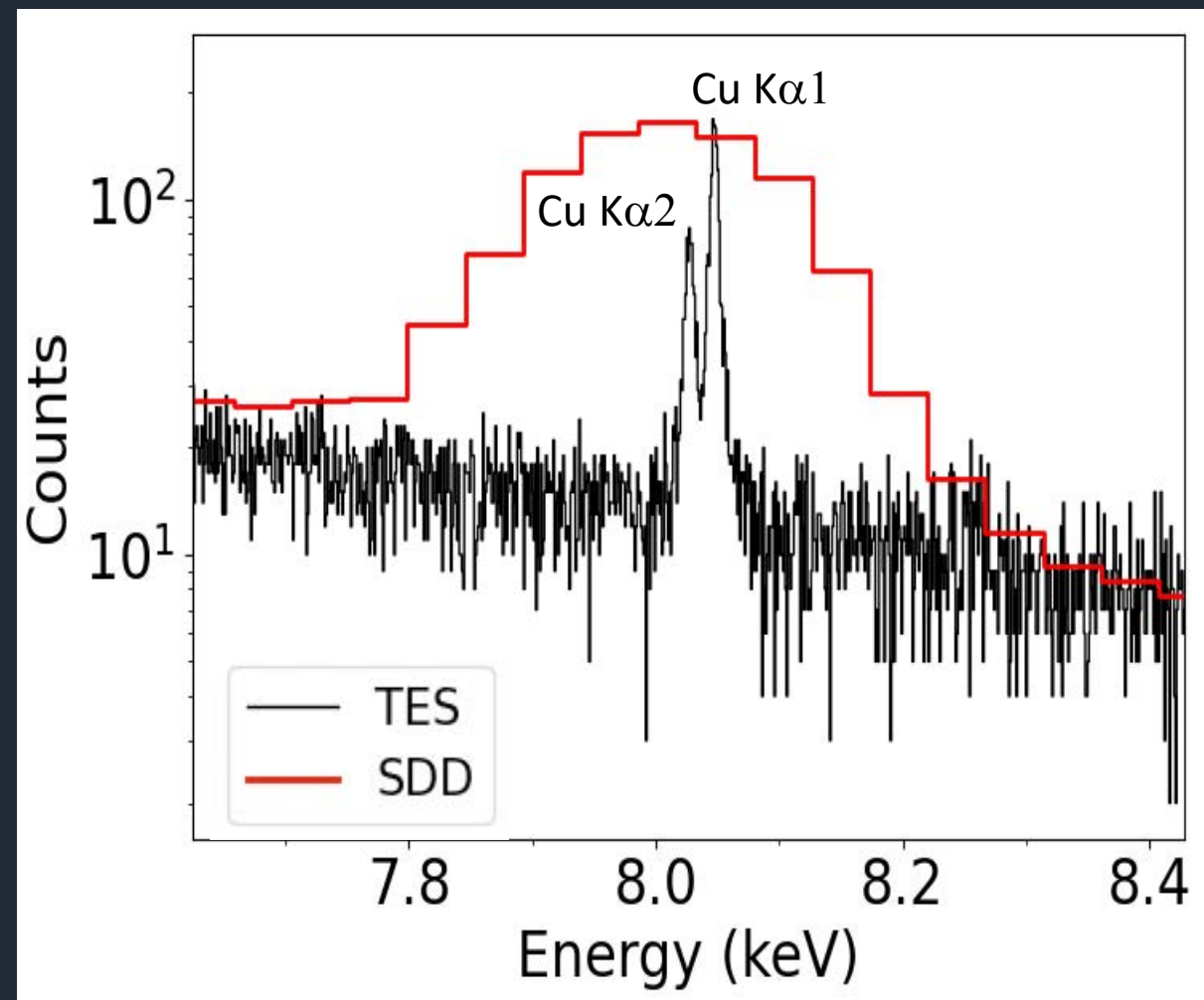
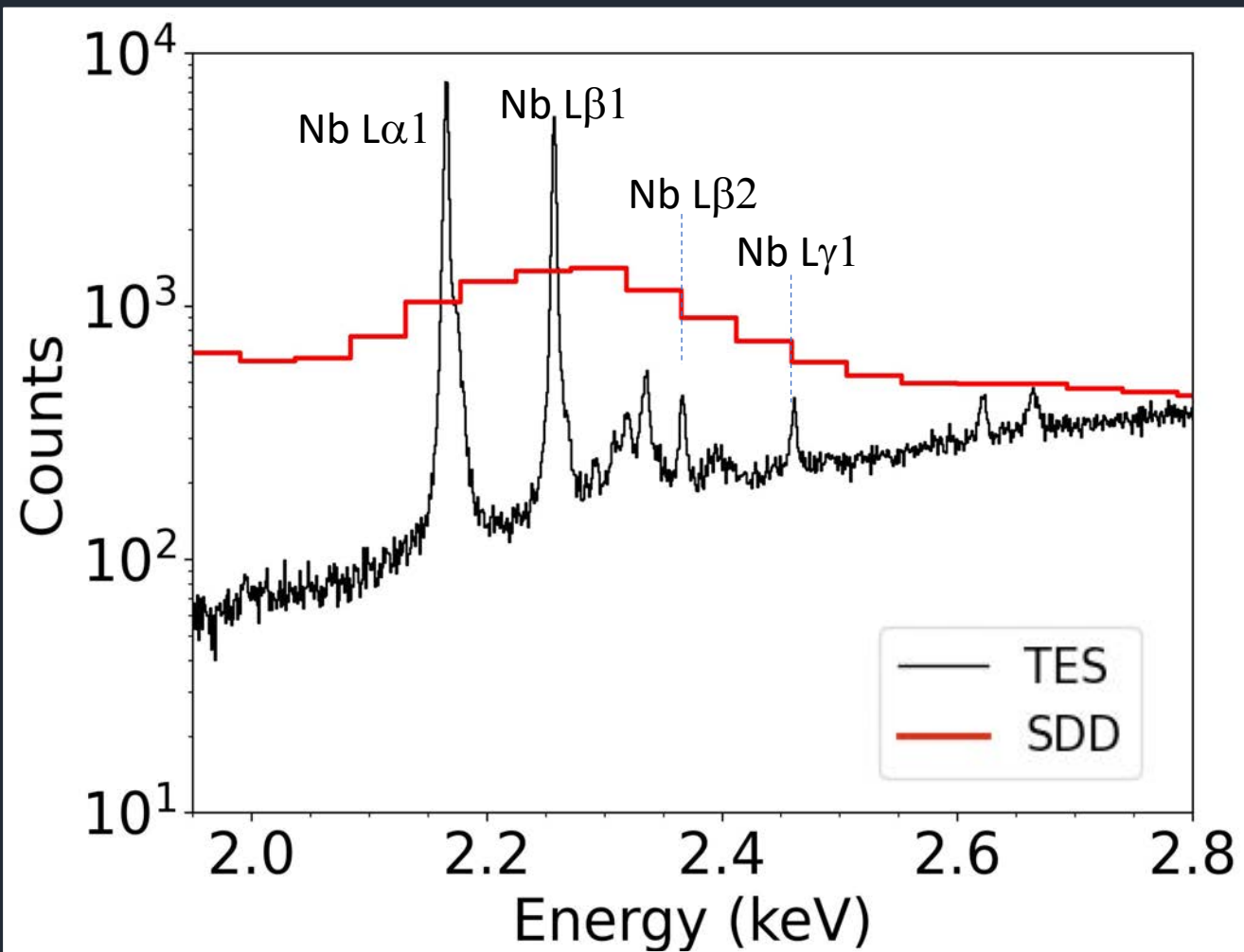
Laboratory high resolution X-ray spectroscopy

- X-ray absorption (XAS),
- X-ray fluorescence (XRF),
- X-ray emission spectroscopy (XES)
- (...)

to measure distribution,
chemical state and local
structure of elements



Laboratory high resolution X-ray spectroscopy

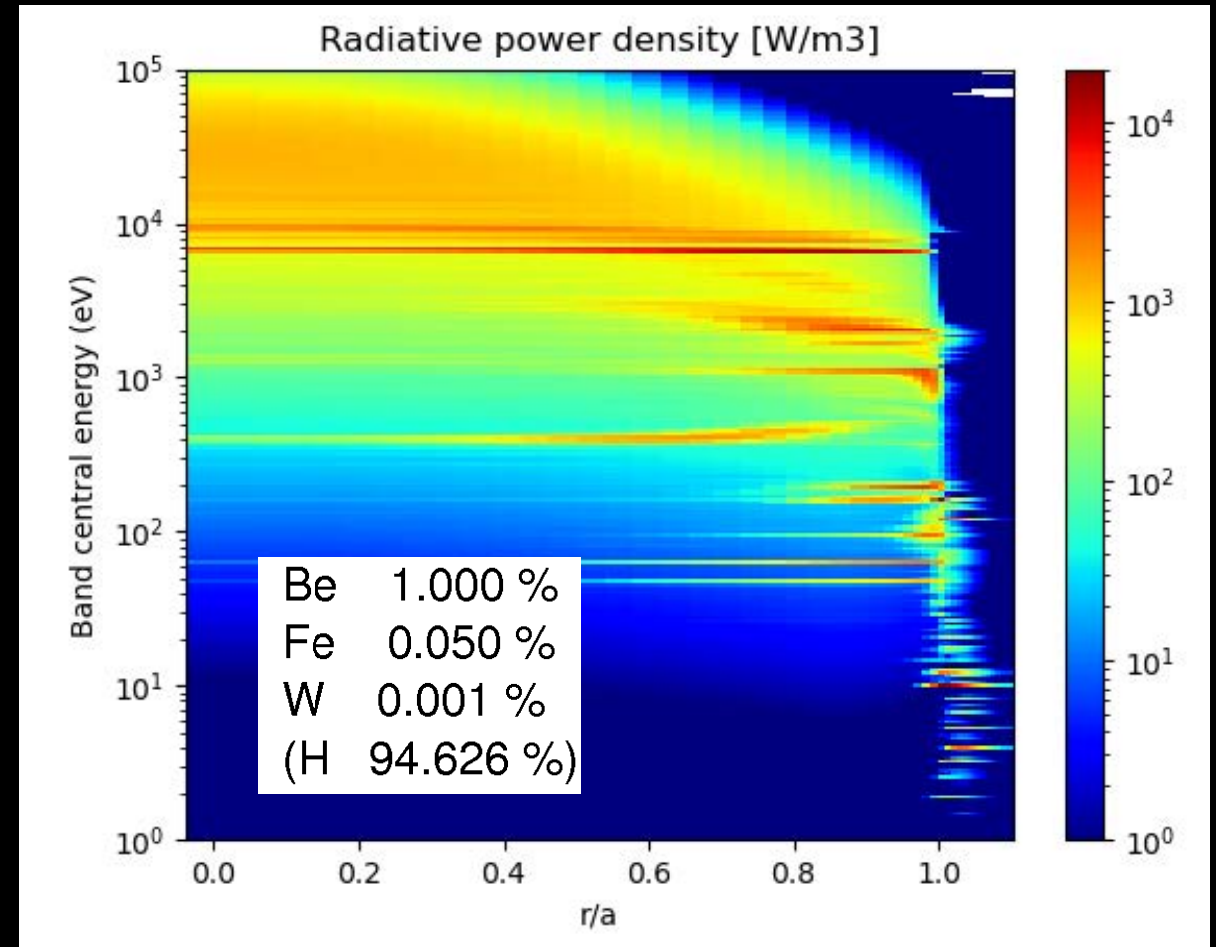


TES X-ray microcalorimeters for fusion plasma physics

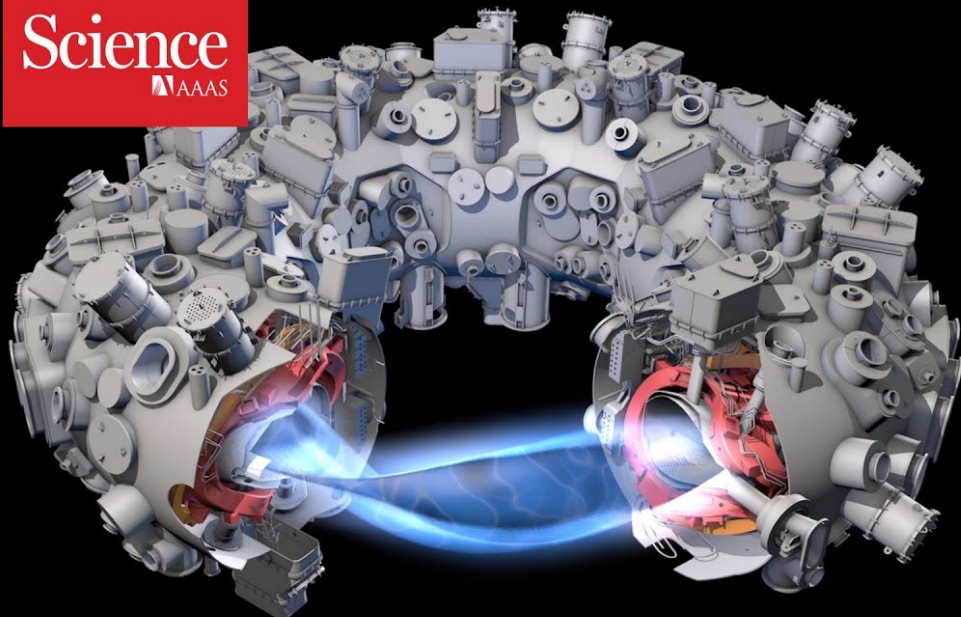
Courtesy: ITER



Tokamak (ITER)



Science
AAAS



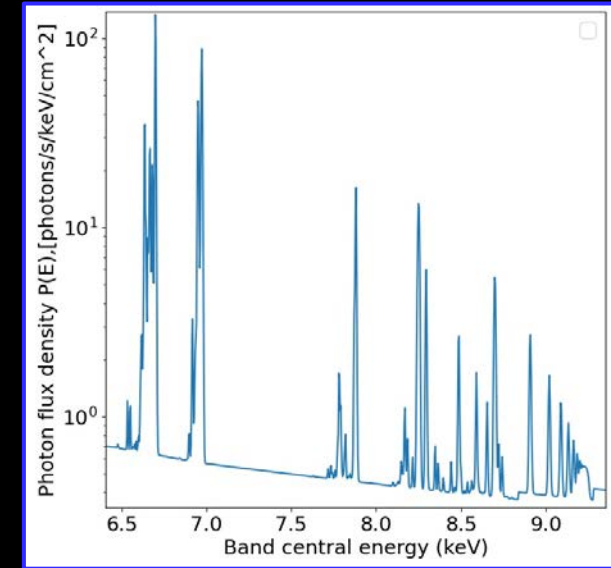
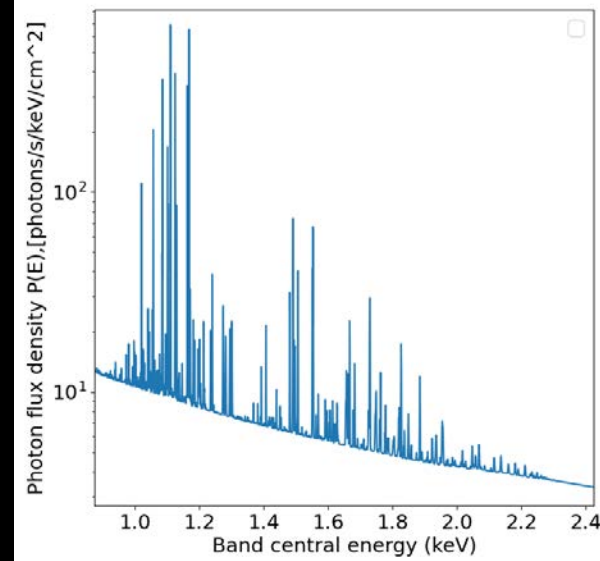
Stellarator (W7-X)

Beiersdorfer et al. RSI 81,10EE323 (2010)
M. Eckart et al. Rev. Sci. Instrum., 92, 063520 (2021)

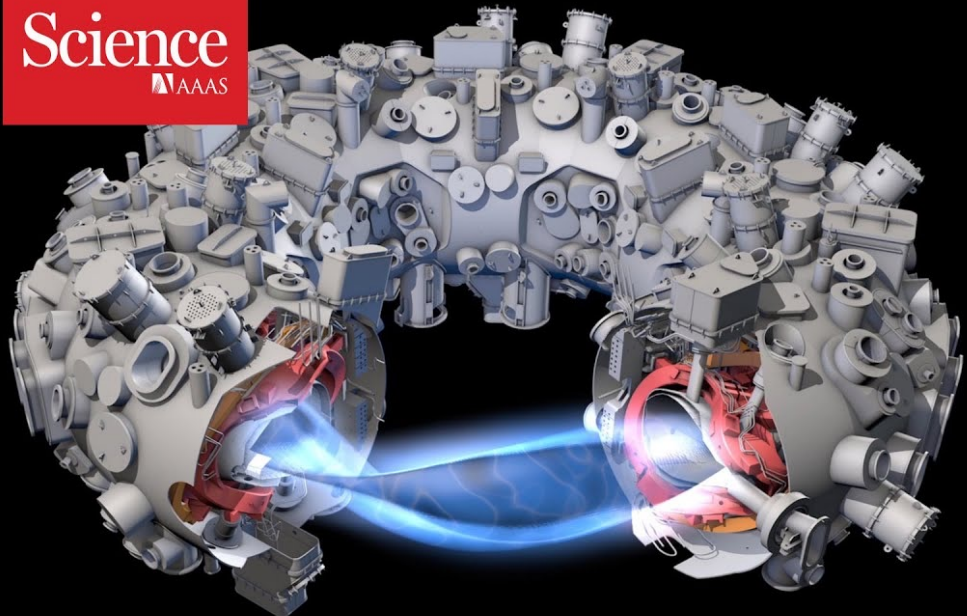
TES X-ray microcalorimeters for fusion plasma physics



Tokamak (ITER)



Science
AAAS



Stellarator (W7-X)

- A XIFU-like X-ray spectrometer could be used for diagnostic purpose in fusion plasmas.

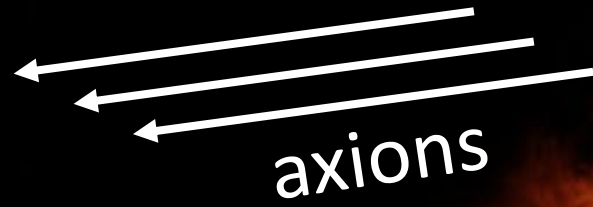
~1600 pixels , $dE \sim 3-5$ eV,
Count-rate capability ~ 1 Mcts/s

Beiersdorfer et al. RSI 81,10EE323 (2010)

M. Eckart et al. Rev. Sci. Instrum., 92, 063520 (2021)

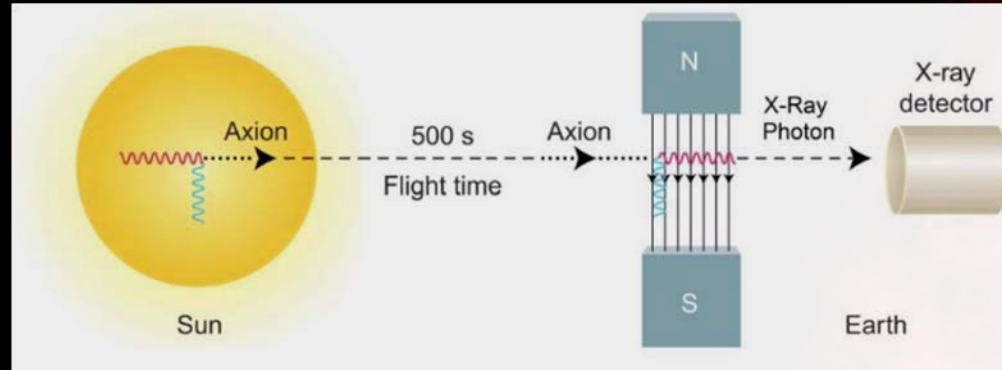
TES X-ray microcalorimeters for solar axion search

Axions telescope



axions

TES X-ray microcalorimeters for solar axion search



R. Battesti et al., Physics Reports 765-766(15), 2018

Axions telescope



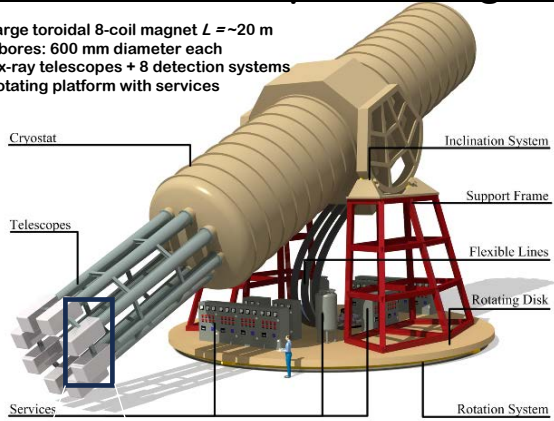
axions





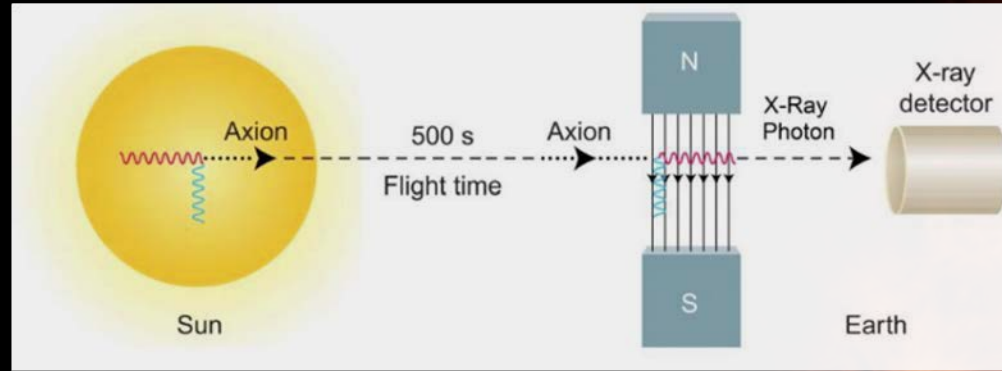
TES X-ray microcalorimeters for solar axion search

- Large toroidal 8-coil magnet $L \approx 20$ m
- 8 bores: 600 mm diameter each
- 8 x-ray telescopes + 8 detection systems
- Rotating platform with services



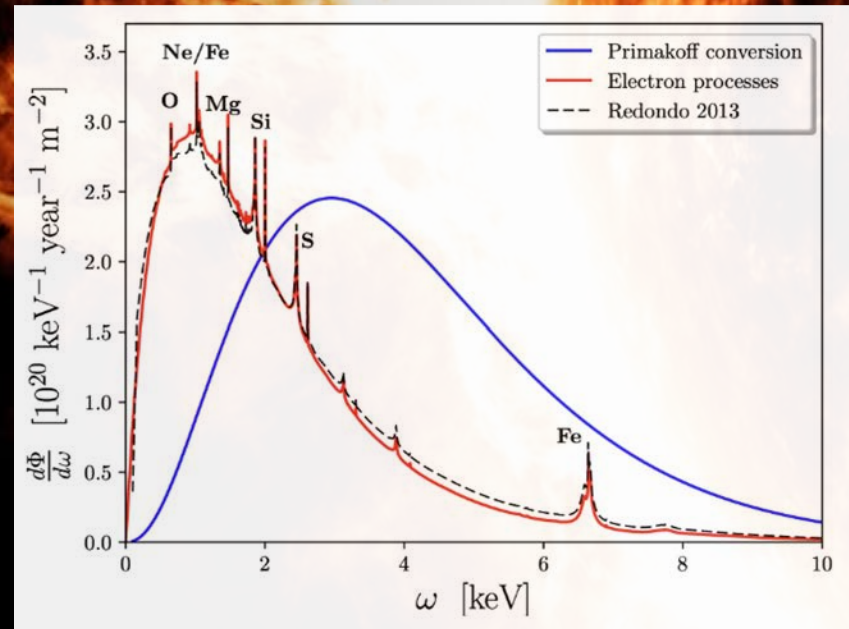
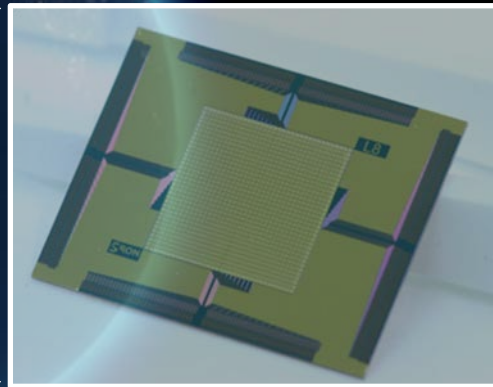
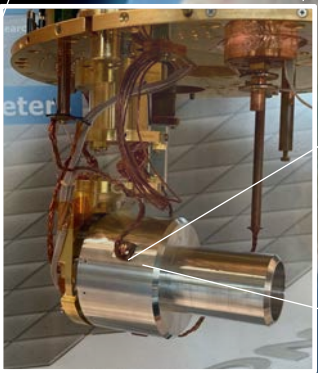
DESY, May 2016
Igor G. Irastorza / Universidad de Zaragoza

<https://faxo.web.cern.ch>



R. Battesti et al., Physics Reports 765-766(15), 2018

axions



Ultra low background TES X-ray spectrometers
D. Vaccaro et al. RSI,94,4 (2023)

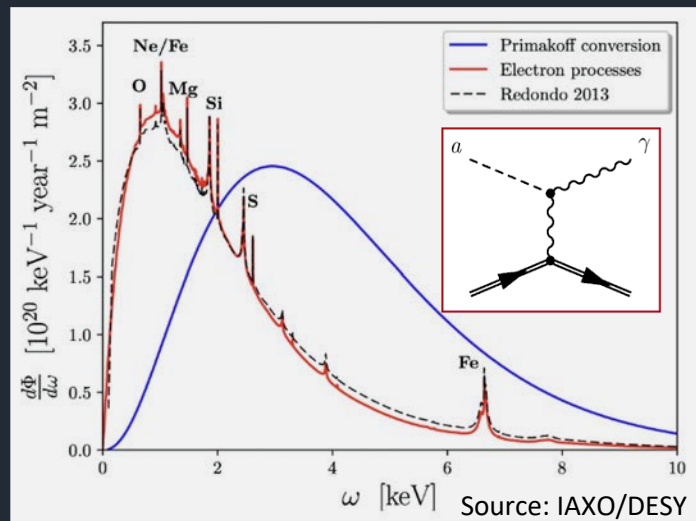
Summary

Superconductivity + Ultra Low Temperature → High-Vision

- X-ray imaging spectrometers based on large array of superconducting transition edge sensors perform at very high resolving power in a wide energy range (0 eV - 20keV)
- They have achieved high level of technological maturity to be used on real instruments
- They can be used in a wide variety of space and ground-based applications

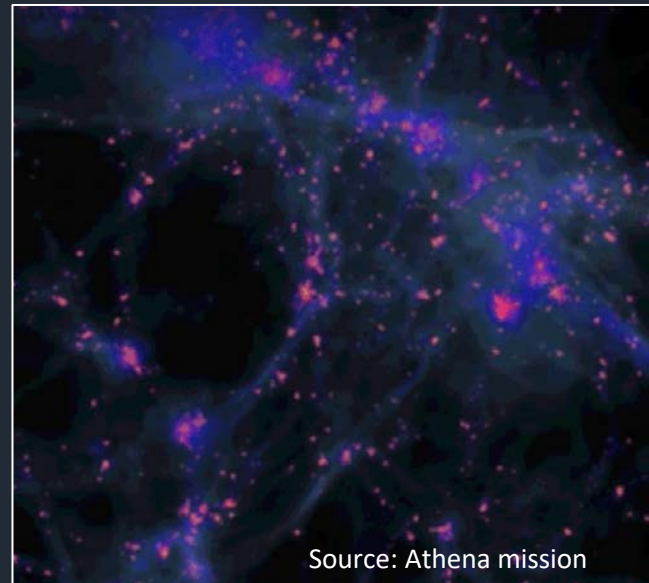
Low background experiments

no photons



X-ray telescopes

some photons



Fusion plasma

many photons



SRON